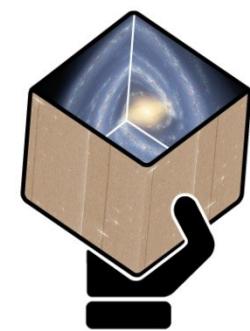


# GENIUS

## Final Report for WP4-T4.4 (INTA)



gaia



*Enrique Solano*  
***Instituto Nacional de Técnica Aeroespacial  
(INTA)***



GENIUS

19 April 2017

# Overview

- **WP4:**
  - **Title:** *Tools for data exploitation*
  - **Goal:** *Development of powerful data exploitation tools allowing a scientific use of the Gaia data beyond the basic queries provided by the main archive interface.*
- **Task 4.4:**
  - **Title:** *VO tools and services*
  - *TOPCAT, VOSA, Clusterix*



# VOSA

<http://svo2.cab.inta-csic.es/theory/vosa/>

Goal: Determination of physical parameters (Teff, logg, [M/H], radius, mass, age,...) from the SED fitting to theoretical models.

File: L Ori-10av.txt (info) (Change)

Model fit+ Template fit

Model fit+ Best fit results

Click in the object name to see the best fits for that object.

Hide graphs Delete this fit Refine excess

Object	RA	DEC	D (pc)	Model	T <sub>eff</sub>	logg	Meta.	more	$\chi^2$	M <sub>d</sub>	F <sub>tot</sub>	ΔF <sub>tot</sub>	F <sub>obs</sub> /F <sub>tot</sub>	L <sub>bol</sub> /L <sub>sun</sub>	ΔL <sub>bol</sub> /L <sub>sun</sub>	λ <sub>max</sub>	A <sub>v</sub>	ΔT <sub>eff</sub>	Δlogg	ΔMeta.	ΔA <sub>v</sub>	N <sub>fit</sub> /N <sub>rot</sub>	Data Vofables
L Ori001	83.446583	9.9273611	400.	BT-Settl CIFIST	3900	4.5	0	alpha:0	9.175e-20	1.389e-20	1.743e-10	2.773e-12	0.59	8.694e-1	1.383e-2	115608	0.1	23.75	0.0865	0	0.0448	17/18	Syn. Spec.
L Ori002	84.043167	10.148583	400.	BT-Settl CIFIST	3900	2.5	0	alpha:0	1.952e+1	1.837e-20	2.417e-10	4.955e-12	0.61	1.205e+0	2.471e-2	115608	0.955	35.71	0	0	0.0547	17/18	Syn. Spec.
L Ori003	83.981000	9.9420833	400.	BT-Settl CIFIST	4300	5.5	0	alpha:0	1.200e+1	1.210e-20	2.223e-10	5.662e-12	0.61	1.108e-0	2.823e-2	115608	0.955	46.73	0.162	0	0.0565	17/18	Syn. Spec.
L Ori004	83.948125	9.7640278	400.	BT-Settl CIFIST	3700	5.5	0	alpha:0	4.849e+0	1.557e-20	1.702e-10	2.535e-12	0.60	8.489e-1	1.265e-2	115608	0.64	0	0.204	0	0.0518	17/18	Syn. Spec.
L Ori005	83.473542	9.7188889	400.	BT-Settl CIFIST	4000	2.5	0	alpha:0	1.110e+1	1.359e-20	1.898e-10	2.243e-12	0.60	9.463e-1	1.118e-2	220883	0.55	9.95	0	0	0.0369	18/18	Syn. Spec.
L Ori006	83.417750	9.9216111	400.	BT-Settl CIFIST	4100	5.5	0	alpha:0	4.541e+0	1.100e-20	1.734e-10	3.221e-12	0.58	8.646e-1	1.606e-2	115608	0.73	26.29	0.0853	0	0.0566	14/15	Syn. Spec.
L Ori007	83.623125	9.8163056	400.	BT-Settl CIFIST	4000	5.5	0	alpha:0	5.181e+0	9.578e-21	1.416e-10	2.838e-12	0.63	7.062e-2	1.415e-2	220883	0.685	33.17	0	0	0.0545	18/18	Syn. Spec.
L Ori008	83.991542	9.9091111	400.	BT-Settl CIFIST	4000	5.5	0	alpha:0	8.610e+0	1.096e-20	1.519e-10	2.867e-12	0.61	7.573e-1	1.430e-2	115608	0.415	19.9	0.0853	0	0.058	17/18	Syn. Spec.
L Ori009	83.693083	10.109889	400.	BT-Settl CIFIST	4300	5	0	alpha:0	1.727e+0	7.100e-21	1.358e-10	1.595e-12	0.63	6.771e-1	7.954e-3	115608	0.685	17.06	0.312	0	0.0276	17/18	Syn. Spec.
L Ori010	83.637333	10.144750	400.	BT-Settl CIFIST	4000	5	0	alpha:0	1.418e+1	7.198e-21	1.042e-10	1.854e-12	0.61	5.198e-1	9.244e-3	220883	0.145	44.9	0.188	0	0.0409	18/18	Syn. Spec.

L Ori001  
BT-Settl-CIFIST, Teff:3900, logg:4.5, Meta.:0, Av:0.1

L Ori002  
BT-Settl-CIFIST, Teff:3900, logg:2.5, Meta.:0, Av:0.955

L Ori003  
BT-Settl-CIFIST, Teff:4300, logg:5.5, Meta.:0, Av:0.955

L Ori004  
BT-Settl-CIFIST, Teff:3700, logg:5.5, Meta.:0, Av:0.64



# VOSA

<http://svo2.cab.inta-csic.es/theory/vosa/>

Available since 2008.  
More than 800 users.  
More than 1.600.000 objects.  
More than 80 refereed papers.

An application to the Collinder 69 open cluster  
A. Bayo<sup>1, 2</sup>, C. Rodríguez Allard<sup>3</sup>

The first planet detected in the WTS: an inflated hot Jupiter in a 3.35 d orbit around a late F star<sup>\*</sup>  
M. Cappella<sup>1, 2</sup>, R. P. Saglia<sup>1, 7</sup>, J. L. Birky<sup>2, 8</sup>, J. Koppenhoefer<sup>2, 9</sup>, D. J. Pinfield<sup>10</sup>, P. Cruz<sup>3</sup>, G. Kovács<sup>3</sup>, B. Sipőcz<sup>3</sup>, D. Barrado<sup>11, 12</sup>, B. Neff<sup>4</sup>, Y. V. Postova<sup>3, 13</sup>, C. del Burgo<sup>10, 11, 12</sup>, E. L. Martin<sup>10</sup>, L. Smiljanic<sup>3</sup>, J. Barnes<sup>5</sup>, A. Bayo<sup>12</sup>, D. A. Gálvez-Ortiz<sup>12</sup>, N. Goulding<sup>5</sup>, C. Haswell<sup>6</sup>, O. Narayan<sup>2</sup>, H. R. Jones<sup>5, 8</sup>, N. Lodieu<sup>5</sup>, F. Marocco<sup>5</sup>, D. Molis<sup>5</sup>, F. Margee<sup>12, 13</sup>, R. Nejowitsch<sup>5</sup>, E. Peña<sup>12</sup>, L. Serio Bero<sup>12</sup>, E. Solano<sup>5, 12</sup>, P. Steele<sup>5</sup>, H. Stover<sup>5</sup>, R. Tata<sup>12, 13</sup> and J. Zerdeja<sup>5</sup>

A&A 550, A144 (2013)

Proper motions of young stars in Chamaeleon II. New kinematical candidate members of Chamaeleon I and II<sup>\*</sup>  
Belén López Martí<sup>1</sup>, Francisco Jiménez-Esteban<sup>1, 2, 3</sup>, Amelia Bayo<sup>4, 5</sup>, David Barrado<sup>1, 6</sup>, Enrique Solano<sup>1, 7</sup>, Hervé Bouy<sup>8</sup> and Carlos Muñoz<sup>1, 2</sup>

Age determination of the HR8799 planetary system using asteroseismology  
A. Moya<sup>1, 2</sup>, P. J. Avendaño<sup>2</sup>, D. Barrado<sup>1, 3</sup>, A. García Hernández<sup>2</sup>, M. Aberasturi<sup>1</sup>, B. Montesinos<sup>1</sup> and A&A 554, A20 (2013)

A Virtual Observatory Census to Address Dwarfs Origins (AVOCADO)  
I. Science goals, sample selection, and analysis tools  
R. Sánchez-Janssen<sup>1</sup>, R. Amorín<sup>2</sup>, M. García-Vargas<sup>3</sup>, J. M. Gómez<sup>4</sup>, M. Huertas-Company<sup>5</sup>, F. Jiménez-Esteban<sup>6, 7, 8</sup>, M. Molina<sup>9</sup>, R. Papaderos<sup>4</sup>, E. Pérez Montero<sup>2</sup>, C. Rodríguez<sup>6, 7</sup>, J. Sánchez Almeida<sup>10, 11</sup> and E. Solano<sup>6, 7</sup>

Searching for transits in the Wide Field Camera Transit Survey with difference-imaging light curves  
J. Zendejas Domínguez<sup>1, 2</sup>, J. Koppenhoefer<sup>2, 1</sup>, R. P. Saglia<sup>2, 1</sup>, J. L. Birky<sup>2</sup>, S. T. Hodgkin<sup>4</sup>, G. Kovács<sup>4</sup>, D. J. Pinfield<sup>2, 5</sup>, B. Sipőcz<sup>3</sup>, D. Barrado<sup>6, 7</sup>, R. Bender<sup>2, 1</sup>, C. del Burgo<sup>6</sup>, M. Cappella<sup>2</sup>, E. L. Martin<sup>9</sup>, S. V. Neff<sup>4</sup>, A. Killeen<sup>1</sup> and P. Steele<sup>2</sup>

A&A 560, A92 (2013)

The Seven Sisters DANCo  
I. Empirical isochrones, luminosity, and mass functions of the Pleiades cluster<sup>\*, \*\*, \*\*\*</sup>  
H. Bouy<sup>1</sup>, E. Bertin<sup>2</sup>, L. M. Sarro<sup>3</sup>, D. Barrado<sup>1</sup>, E. Moraux<sup>4</sup>, J. Bourde<sup>4</sup> and V. Belotsky<sup>7</sup>

A&A 574, A57 (2015)

The CoRoT chemical peculiar target star HD 49310<sup>\*</sup>  
J. Gilliland<sup>1</sup>, W. W. Weiss<sup>3</sup> and T. Lüftinger<sup>3</sup>

The Astrophysical Journal Supplement Series > Volume 236 > Number 2  
A GALEX-based Search for the Sparse Young Taurus-Aurigae Star Forming Region  
Ana I. Gómez de Castro<sup>1</sup>, Javier López-Santiago<sup>1</sup>, Fatima López-Martínez<sup>2</sup>, Néstor Sánchez<sup>3</sup>, Paolo Manolli Comide<sup>4</sup> and Javier Yáñez Gestoso<sup>4</sup>

A&A 558, A116 (2013)

HD 85567: A Herbig B[e] star or an interacting B[e] binary?  
Resolving HD 85567's circumstellar environment with the VLT and AMBER<sup>\*, \*\*</sup>  
H. E. Wheelwright<sup>1</sup>, G. Weigelt<sup>1</sup>, A. I. Gómez de Castro<sup>2</sup>, J. L. Alcalá<sup>3</sup>, J. L. Fernández-Prado<sup>4</sup>, J. L. Gómez<sup>5</sup>, A. L. Rodriguez<sup>6</sup>, A. Mennickent<sup>1, \*</sup>, G. Djurasevic<sup>7</sup>, M. Cabezas<sup>1</sup>, A. Codzik<sup>7</sup>, J. G. Rosales<sup>7</sup>, E. Niemczura<sup>7</sup>, I. Anaya<sup>7</sup> and M. Cunha<sup>7</sup>

A&A 541, A38 (2012)

High-resolution imaging of Kepler planet host candidates: A comprehensive comparison of different techniques<sup>\*</sup>  
J. Lillo-Box<sup>1</sup>, D. Barrado and H. Bouy<sup>2</sup>

A&A 541, A38 (2012)

Fundamental parameters of the close interacting binary HD 170582 and its luminous accretion disc<sup>\*</sup>  
R. E. Mennickent<sup>1, \*</sup>, G. Djurasevic<sup>2</sup>, M. Cabezas<sup>1</sup>, A. Codzik<sup>2</sup>, J. G. Rosales<sup>2</sup>, E. Niemczura<sup>2</sup>, I. Anaya<sup>2</sup> and M. Cunha<sup>2</sup>

A&A 550, A163 (2014)

Warm debris disks candidates in transiting planets systems  
A. Ribas<sup>1</sup>, B. Merin<sup>1</sup>, D. R. Ardila<sup>2</sup> and H. Bouy<sup>1</sup>

A&A 541, A38 (2012)



# VOSA

<http://svo2.cab.inta-csic.es/theory/vosa/>

## SED building: Gaia DR1 included in the list of photometric catalogues VOSA uses to build the SED.



# VOSA

<http://svo2.cab.inta-csic.es/theory/vosa/>

Phys. parameter determination: TGAS distances available from VOSA to estimate bolometric luminosities.



GENIUS

19 April 2017



# VOSA

<http://svo2.cab.inta-csic.es/theory/vosa/>

## New architecture to scale VOSA to Big Data:

Distributed environment.

Parallelized computing.

Asynchronous jobs.

Front-end redesigned.

**VOSA is now able to handle files with tens of thousands of objects.**

Files      Objects      Build SEDs      **Analyse SEDs**      HR Diag.      Results      Help

Test: Stars and brown dwarfs ([Change](#))      File: jplus2 ([info](#)) ([Change](#))

[Model Fit](#)      Template fit      Model Bayes Analysis      Template Bayes Analysis      Binary fit

**Model fit**

The fit process has been submitted asynchronously.

6.73% completed

I expect the whole process to take around 34 minutes to finish  
Please, refresh this page again in a while for updated info

Clusterix 2.0 is an interactive web-based application to calculate the grouping probability of a list of objects using proper motions and the non parametric method described in Galadi-Enriquez et al. 1998. It also allows the possibility of gathering physical parameters (parallaxes, radial velocities, proper motions,...) from Vizier and estimating effective temperatures, surface gravities and metallicities using VOSA.

## Step 1/3: Information gathering (coordinates and physical parameters)

<b>Search by Id</b> Usage: ID M45 Radius 5 deg Catalogue TGAS/DR1 <input type="button" value="Search"/>	<b>Search by Coordinates</b> Usage: RAJ2000(deg),DEJ2000(deg) <input type="text"/> Radius arcmin Catalogue GAIA/DR1 <input type="button" value="Search"/>
<b>Search in Webda ?</b> <input type="button" value=""/>	<b>Search by file?</b> <input type="button" value="Choose file"/> No file chosen
<input type="button" value="Clear"/>	

Clusterix 2.0 is the result of a collaboration agreement between the Centro de Astrobiología (CAB, INTA-CSIC) and the Universidad de Barcelona based on a previous work done at the Masaryk University. It has been developed in the framework of the GENIUS project and is maintained by the Spanish Virtual Observatory at the Data Archive Unit of the CAB (INTA-CSIC). If you use this service in your research, please include the following acknowledgment in any resulting publications: "Based on Clusterix 2.0 service at CAB (INTA-CSIC)".

This project has received funding from the European Union's Seventh Framework Programme (FP7-SPACE-2013-1) for research, technological development and demonstration under grant agreement no. 606740.

SVO - CAB (INTA-CSIC). Version 1.0



## Step 2/3: Region selection

Cluster info: M45\_5\_deg\_TGAS

### Selection of the "cluster" and "field" regions

Area type:  Polygon  Circle

Cluster (shift):

56.9489,23.9631,1.0275;

Field (ctrl):

59.2397,21.7632 57.1368,2

### Membership determination parameters

Maximum  $\mu$  (mas/yr):

10 $\mu$ .0

Maximum  $\mu$  err (mas/yr):

10.0

Smooth param (no unit):

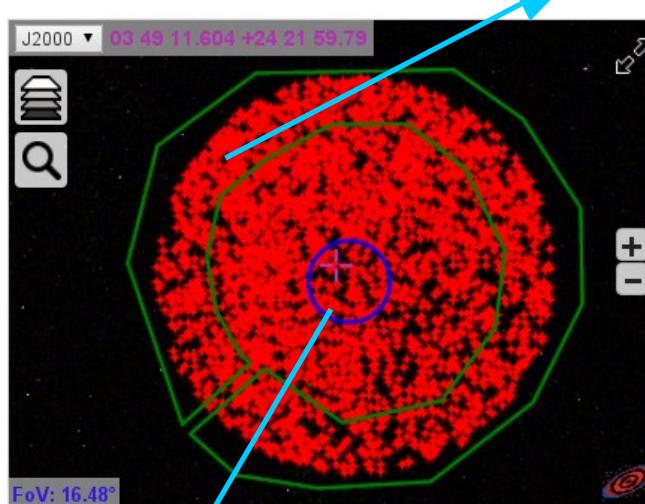
6.420576433

Silverman rule: 6.420576433

$\gamma$  threshold:

3.0

Submit Clean



$$\Psi_f^k(\mu_{xi}, \mu_{yj})$$

$$\Psi_{c+f}^k(\mu_{xi}, \mu_{yj})$$

To draw areas select the area type (polygon or circle) in the radio button and press **left-shift** key for defining a cluster and **ctrl** key for defining a field.

clicking on the image. After the second point you will see the **shift/ctrl** and click on the image. That will close the polygon. center of the polygon.

the center.

Release the **shift/ctrl** key and click on the desired outer limit/radius of the circle you want to draw. The blue circle will disappear and the final circle will appear.

Coordinates must be in J2000 (deg)



GENIUS

19 April 2017

## Step 3/3. Determination of membership probabilities

SAMP

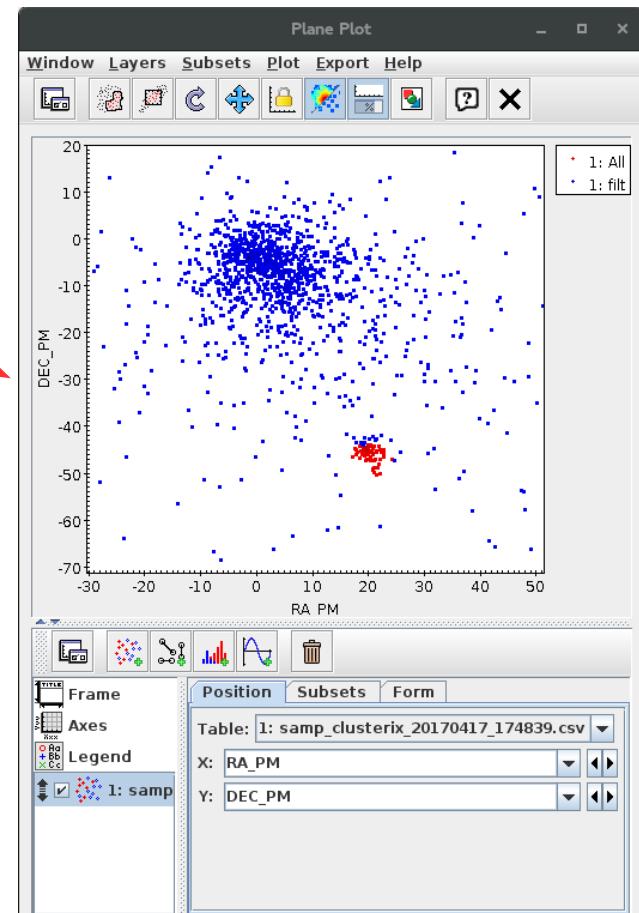
```

# Results were retrieved using Clusterix software
# http://clusterix.cab.inta-csic.es/
# In case of problems, please, report to: clusterix_archive_support@cab.inta-csic.es
#
# Labels:
#
# STAR_NO identifier of star retrieved from the input data
# RA right ascension of a star
# DEC declination of a star
# PM_RA proper motion in alpha
# PM_DEC proper motion in delta
# PROB probability that star belongs to evaluated open cluster
# FLAG M=Membership NM=No membership
#
# Parameters:
#
# CLUSTER INFO: M45_5_deg_TGAS
# PROPER MOTION CUTOFF: 100.00 mas/yr
# PROPER MOTION ERR CUTOFF: 10.00 mas/yr
# SMOOTH PARAMETER: 6.42
# GAMMA FACTOR: 3.00
#
# Boundaries can be circles (Ra(J2000)-center,Dec(J2000)-center, radius degrees) or
# polygons (Ra,Dec Ra,Dec ...). Each boundary is separated by ;
#
# CLUSTER BOUNDARIES:
# 56.9489,23.9631,1.0275;
#
# FIELD BOUNDARIES:
# 59.2397,21.7632 57.1368,20.4117 54.4734,20.9335 52.9596,22.7285 52.6240,24.6696 53.6630,26.7712 55.2726,27.2971 59.6181,29.1955 53.9118,29.2663 51.9734,27.9665 50.4601,25.6624 50.5435,23.2572 51.9978,20.8227 54.786
#
# EXPECTED NUMBER OF MEMBERS: 61
#
# For Webda clusters, STAR_NO is the merge of the Data source reference and the star number: Ref_Star
#
#STAR_NO      RA (J2000)      DEC (J2000)      PM_RA      PM_DEC      PROB FLAG    Initial guess
117          57.820366     23.826412     20.27     -47.40   0.97    M       C
152          57.987022     23.901766     19.73     -47.15   0.97    M       C
11           56.657089     23.787523     19.65     -47.11   0.97    M       C
162          57.714436     23.328882     19.51     -47.28   0.97    M       C

```

Save as

$$P_c^k(\mu_{xi}, \mu_{yj}) = \frac{\Psi_{c+f}^k(\mu_{xi}, \mu_{yj}) - \Psi_f^k(\mu_{xi}, \mu_{yj})}{\Psi_{c+f}^k(\mu_{xi}, \mu_{yj})}$$



## Gather information from VO services

Proper Motion

Radius

arcsec ▾

TGAS ▾

Radial Velocity

Radius

arcsec ▾

Gaia ESO ▾

Parallax

Radius

arcsec ▾

TGAS ▾

VOSA Photometry ?

Clusterix 2.0 is the result of a collaboration agreement between the Centro de Astrobiología (CAB, INTA-CSIC) and the Universidad de Barcelona based on a previous work done at the Masaryk University. It has been developed in the framework of the GENIUS project and is maintained by the Spanish Virtual Observatory at the Data Archive Unit of the CAB (INTA-CSIC)

If you use this service in your research, please include the following acknowledgment in any resulting publications: "Based on Clusterix 2.0 service at CAB (INTA-CSIC)".

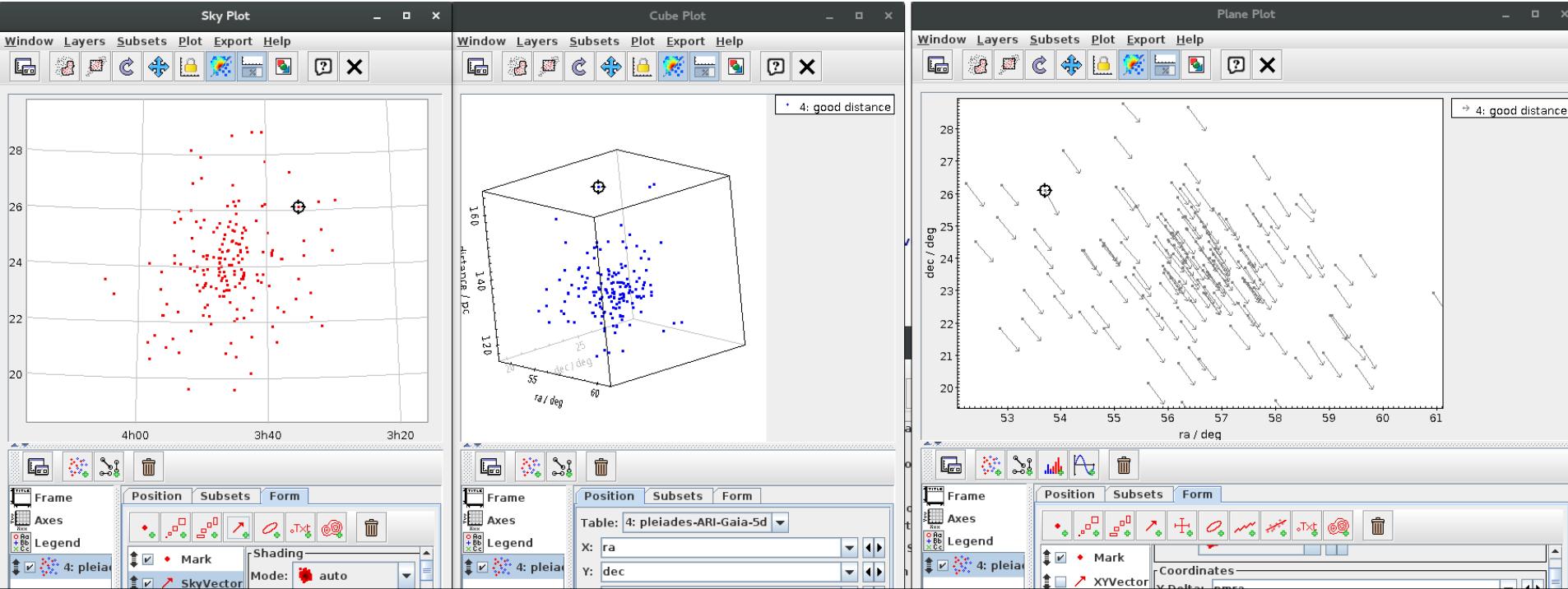
This project has received funding from the European Union's Seventh Framework Programme (FP7-SPACE-2013-1) for research, technological development and demonstration under grant agreement no. 606740.

SVO - CAB (INTA-CSIC). Version 1.0



# Clusterix 2.0

<http://clusterix.cab.inta-csic.es>



GENIUS

19 April 2017

# Sustainability

<http://clusterix.cab.inta-CSIC.es>

- *VOSA and Clusterix have been developed in the framework of the Spanish Virtual Observatory, a project hosted at INTA since 2004.*
  - *Maintenance and new developments (on a best effort basis) are guaranteed.*

## Future enhancements:



Tangential velocities instead of proper motions.

Access to Gaia DR2 data.

N-dimensional scenario.

Asynchronous jobs.