



# CARMENES



## $\sigma$ Orionis



## Gaia DR2

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Vicky



Cristina



Barcelona

**Woody Allen**

Bardem, Cruz, Johansson, Hall...

# BajaMasa RecGaia research lines



- **EXOS:** exoplanetary systems
  - EXOS-1: astrometry of known systems
  - EXOS-2: radial velocity of new systems
  - EXOS-3: detailed characterisation
- **MLT:** ultracool dwarfs
  - MLT-1: late M (H-R diagrams, kinematics...)
  - MLT-2: L and T (isolated or companions)
- **YBD:** young brown dwarfs
  - Bottom of the (I)MF in young open clusters and stellar associations

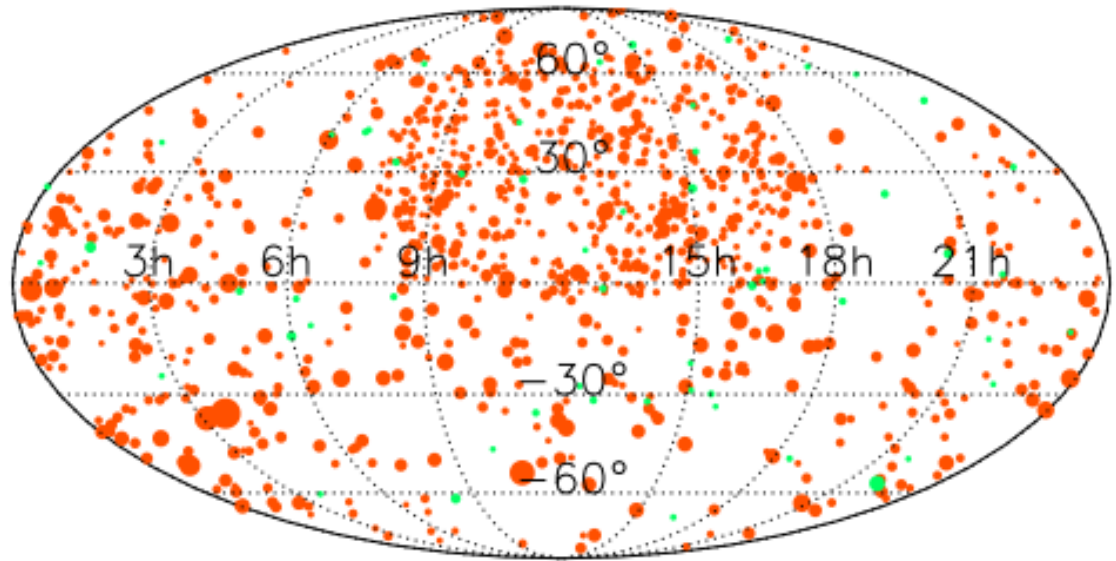
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# Gaia & BDs 2017

The Gaia ultracool dwarf sample – I. Known L and T dwarfs in the first Gaia data release  
Smart, Marocco,  
Caballero et al.  
2017MNRAS.469.401S

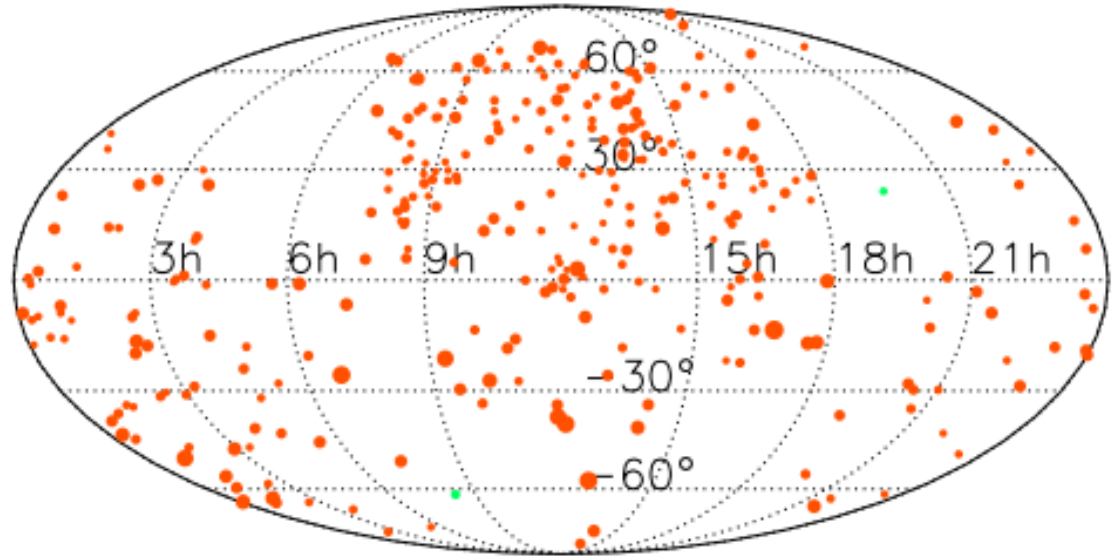


**Figure 4.** The equatorial distribution of 1010 L (red) and 58 T (green) dwarfs with  $G_{\text{est}} < 21.5$  mag. The size of the symbol indicates the  $G_{\text{est}}$  magnitude – larger is brighter.



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Smart, Marocco,  
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2017MNRAS.469.401S

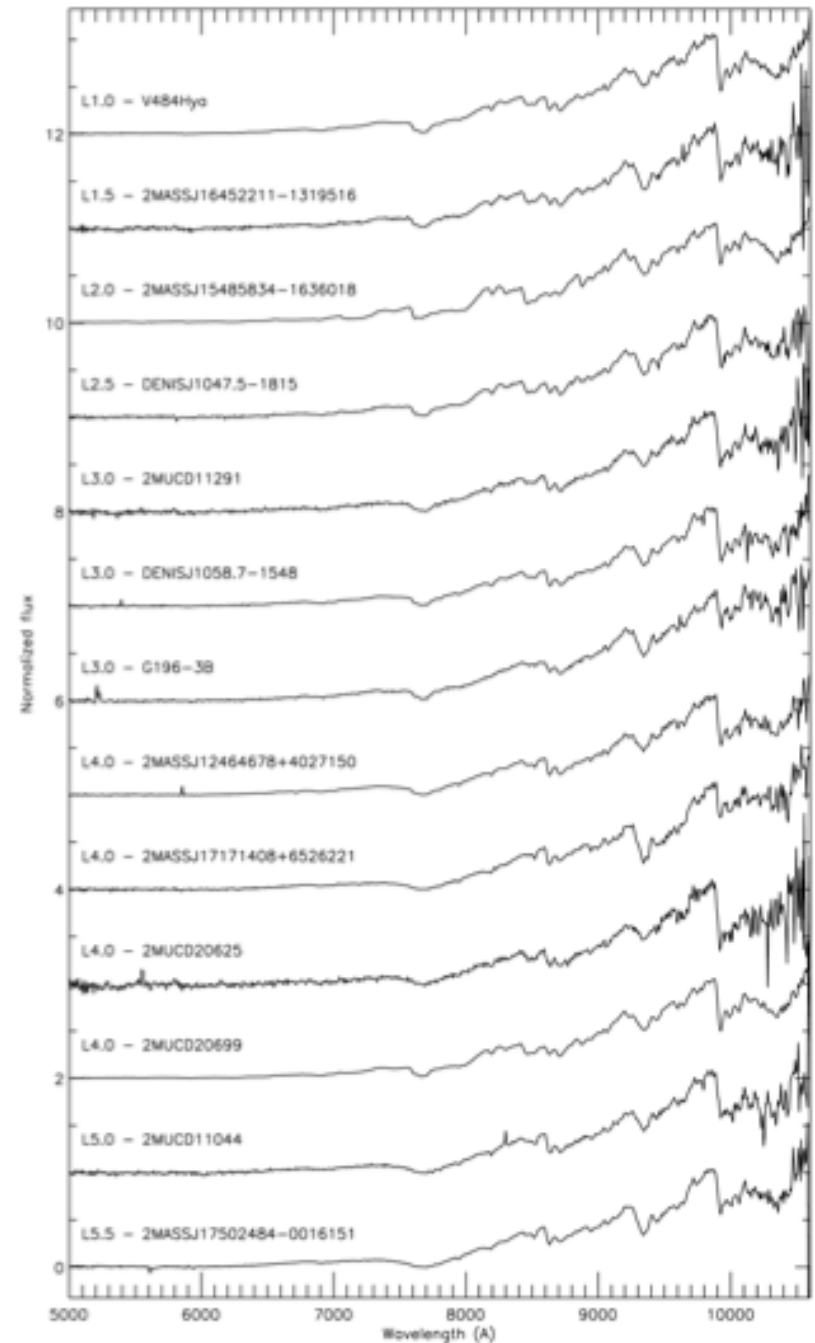


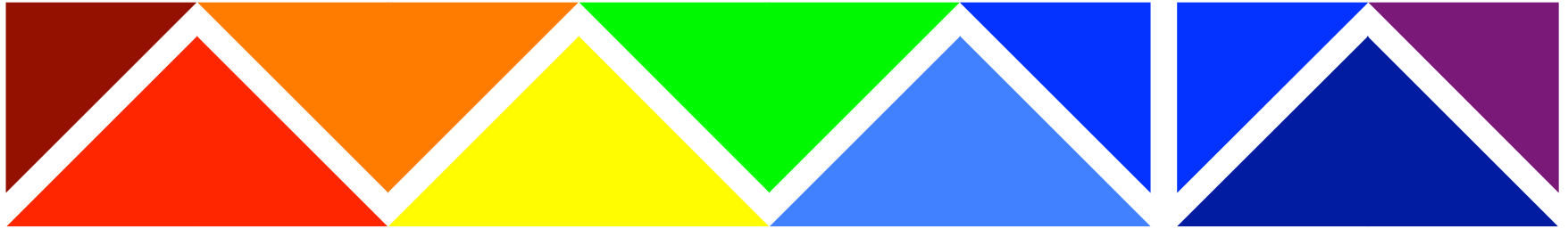
**Figure 6.** Same as Fig. 4, but for the 321 L (red) and T (green) dwarfs with an entry in *Gaia* DR1.



# Gaia & BDs 2018

The Gaia ultracool dwarf sample – II. Known L and T dwarfs in the **second** Gaia data release  
Smart et al.  
2018, MNRAS  
to be submitted





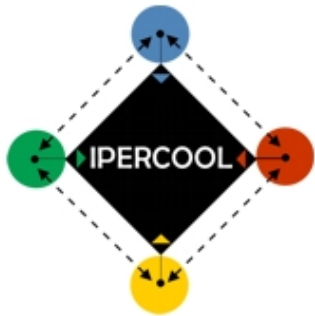
M-, L- AND T-DWARF ARCHIVE OF INTEREST FOR ASTROPHYSICS

# MAIA

*Gaia*, brown dwarfs and cool neighbours: an **M-, L- and T-dwarf Archive of Interest for Astrophysics**







# IPERCOOL

Interpretation and Parameterization of Extremely Red COOL dwarfs

HOME . PARTNERS . WORK PLAN . PARAMETERS . INTERPRETATION . DATA RELEASES . PUBLICATIONS .

INTERNAL

HOME



## Project summary

The **IPERCOOL** project combines the expertise of Torino Observatory (INAF-OATo), the Center for Astrophysical Research at the University of Hertfordshire (CAR-UH), the National Brazilian Observatory (ON/MCT) and the Shanghai Astronomical Observatory (SHAO) to the observation and scientific interpretation of low mass star and brown dwarfs. We pursue active observational and compilational programs that measure the distance, colors and spectra of these objects to find their physical parameters, test atmospheric models, determine parameters of the Galaxy and the history of brown dwarf formation.

The heart of the project is a **International Research Staff Exchange Scheme** funded by the European Union via the **IPERCOOL # 247593** grant within the Marie Curie 7th European Community Framework Programme.

The logo illustrates the overall structure: exchanges are made only between the European and Non-European partners and all of the consortium can share data via a central mysql DB with web and direct interfaces.



This project is funded by the European Union via the International Research Staff Exchange Scheme within the Marie Curie 7th European Community Framework Programme.

IPERCOOL → MAIA



# RESEARCH & INNOVATION

European Commission

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COSME

3rd Health Programme

Consumer Programme

### FP7 & CIP Programmes 2007-2013

Calls

Other Funding Opportunities

## MARIE SKŁODOWSKA-CURIE RESEARCH AND INNOVATION STAFF EXCHANGE (RISE)

H2020-MSCA-RISE-2015

|                     |             |                      |   |
|---------------------|-------------|----------------------|---|
| <b>Opening Date</b> | 06-01-2015  | <b>Deadline Date</b> | 28-04-2015 17:00:00 (Brussels local time)   |
| <b>Budget</b>       | €80,000,000 | <b>Programme</b>     | Horizon 2020                                |
| <b>Status</b>       | <b>Open</b> | <b>Main Pillar</b>   | Excellent Science                           |
|                     |             | <b>OJ reference</b>  | <a href="#">OJ C361 of 11 December 2013</a> |

Call description

[Call documents](#)

[Get support](#)

### Call updates

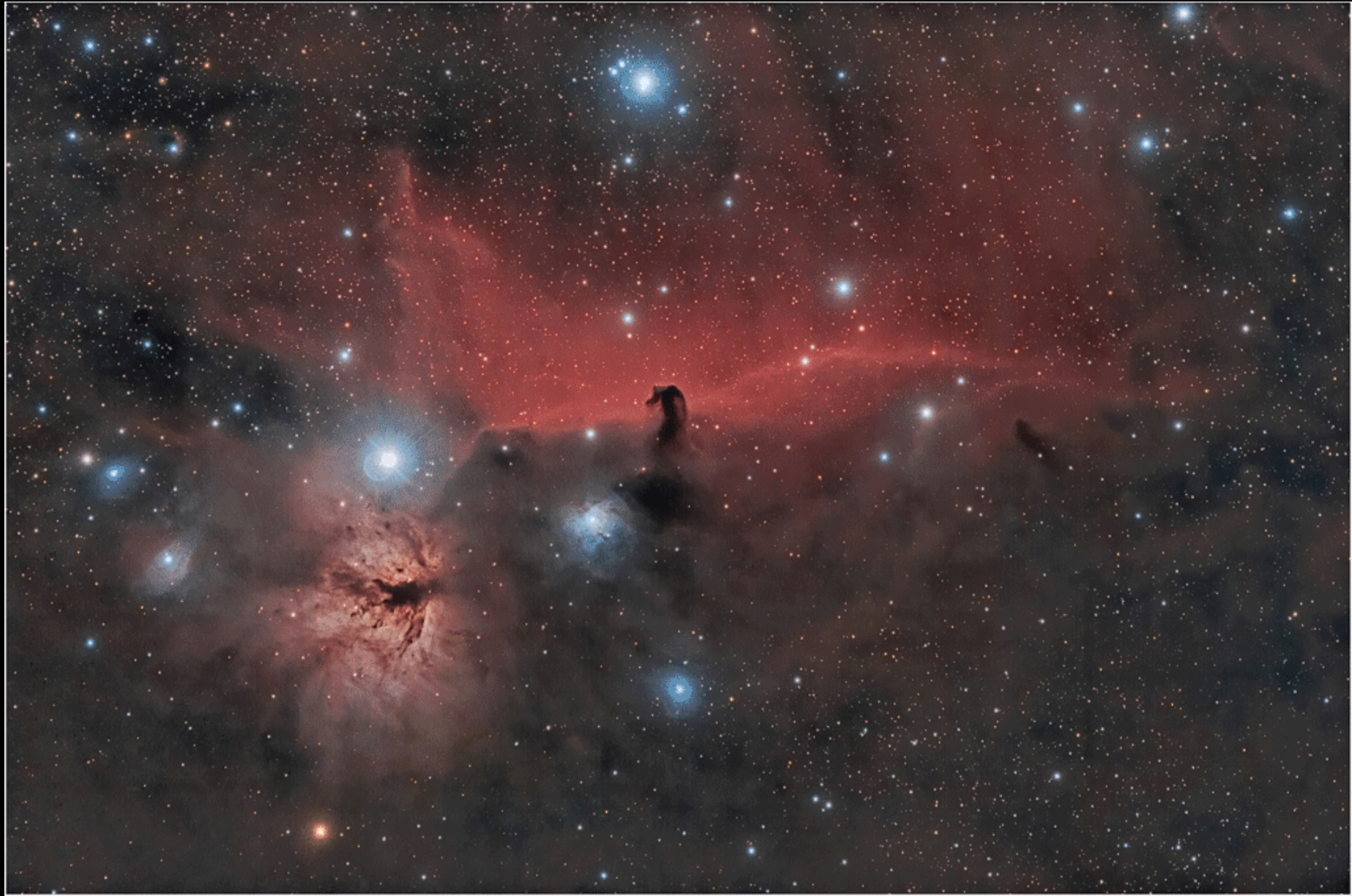
• **13-03-2015 17:02:37**

A revised RISE 2015 - FAQs (under [Call documents](#)) has been updated and published.

[+ More](#)

Topics and submission service

*The East of the Orion Belt (Amitak,  $\sigma$  Orionis, The Horsehead, The Flame...)*



*Raúl Acaraz Gómez, Astronomía A.C. Caballero et al.*

# $\sigma$ Orionis: basic parameters

Ori OB1b  
(or not?)

Age  $\tau \sim 3$  Ma  
(2-4 Ma)

Distance  $d \sim 385$  pc  
(352-440 pc)

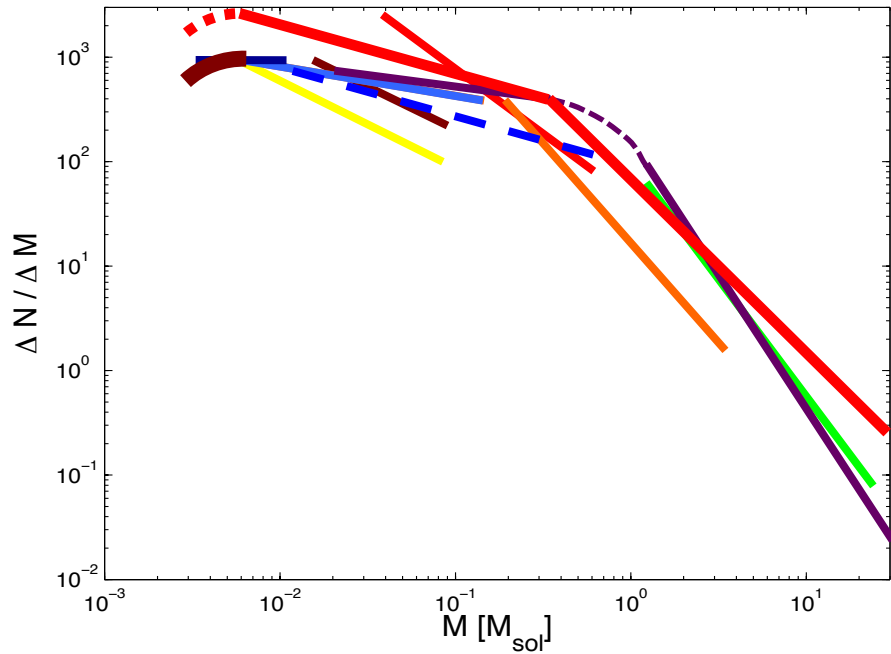
Radius  $\rho \sim 30$  arcmin  
(20-arcmin core)

Extinction  $A_V < 0.3$  mag  
(very low  $E(B-V)$ )

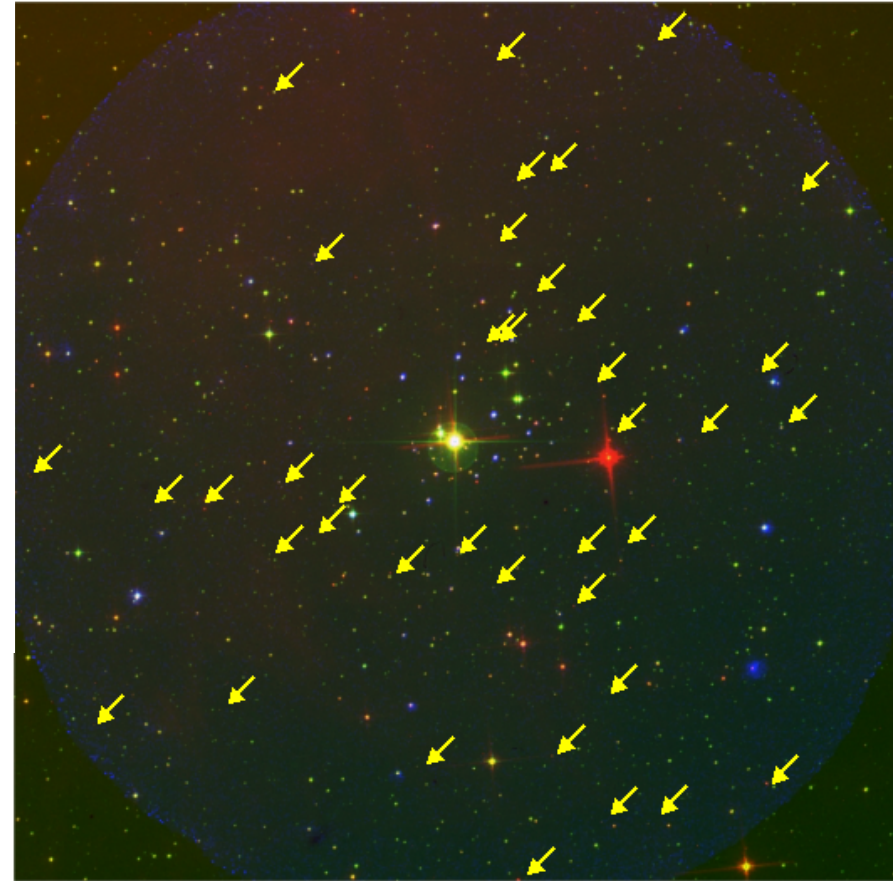
Solar metallicity  
( $[Fe/H] = 0.0$ )



# The $\sigma$ Orionis (initial) mass function



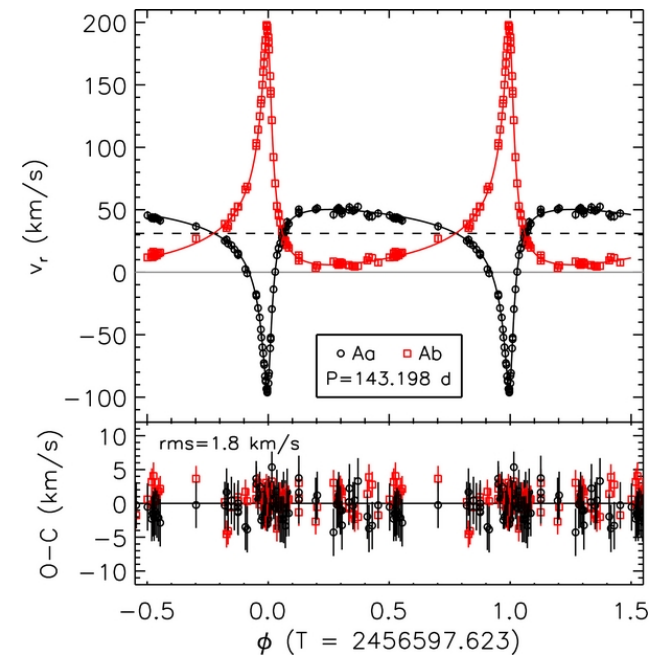
**Distance**  $d \sim 385$  pc (?)  
**Proper motion**  $\mu < 10$  mas  $\text{a}^{-1}$





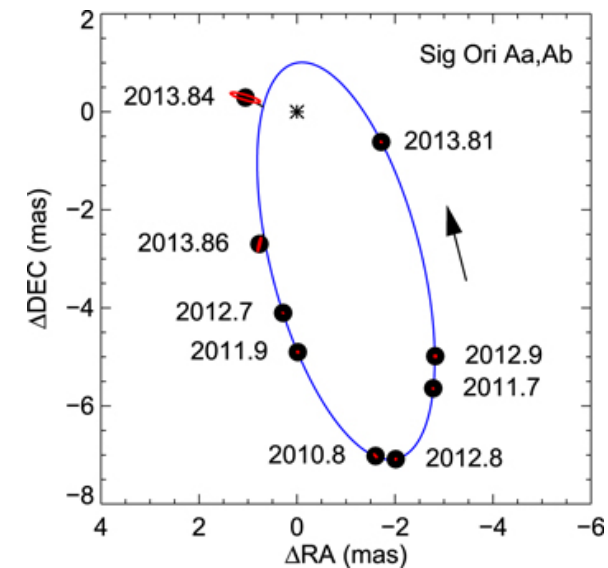
# A precise (and accurate?) distance to $\sigma$ Ori

- Simón-Díaz, Caballero et al. (2011, 2015)  $\rightarrow$  Hi-res spectroscopy
- Schaefer et al. (2016)  $\rightarrow$  Optical interferometry
- $d = 387.5 \pm 1.3$  pc



**Table 11**  
Derived Properties for  $\sigma$  Orionis Aa, Ab, and B

| Parameter                | Value               |
|--------------------------|---------------------|
| $M_{Aa}$ ( $M_{\odot}$ ) | $16.99 \pm 0.20$    |
| $M_{Ab}$ ( $M_{\odot}$ ) | $12.81 \pm 0.18$    |
| $M_B$ ( $M_{\odot}$ )    | $11.54 \pm 1.15$    |
| $\pi$ (mas)              | $2.5806 \pm 0.0088$ |
| $d$ (pc)                 | $387.51 \pm 1.32$   |

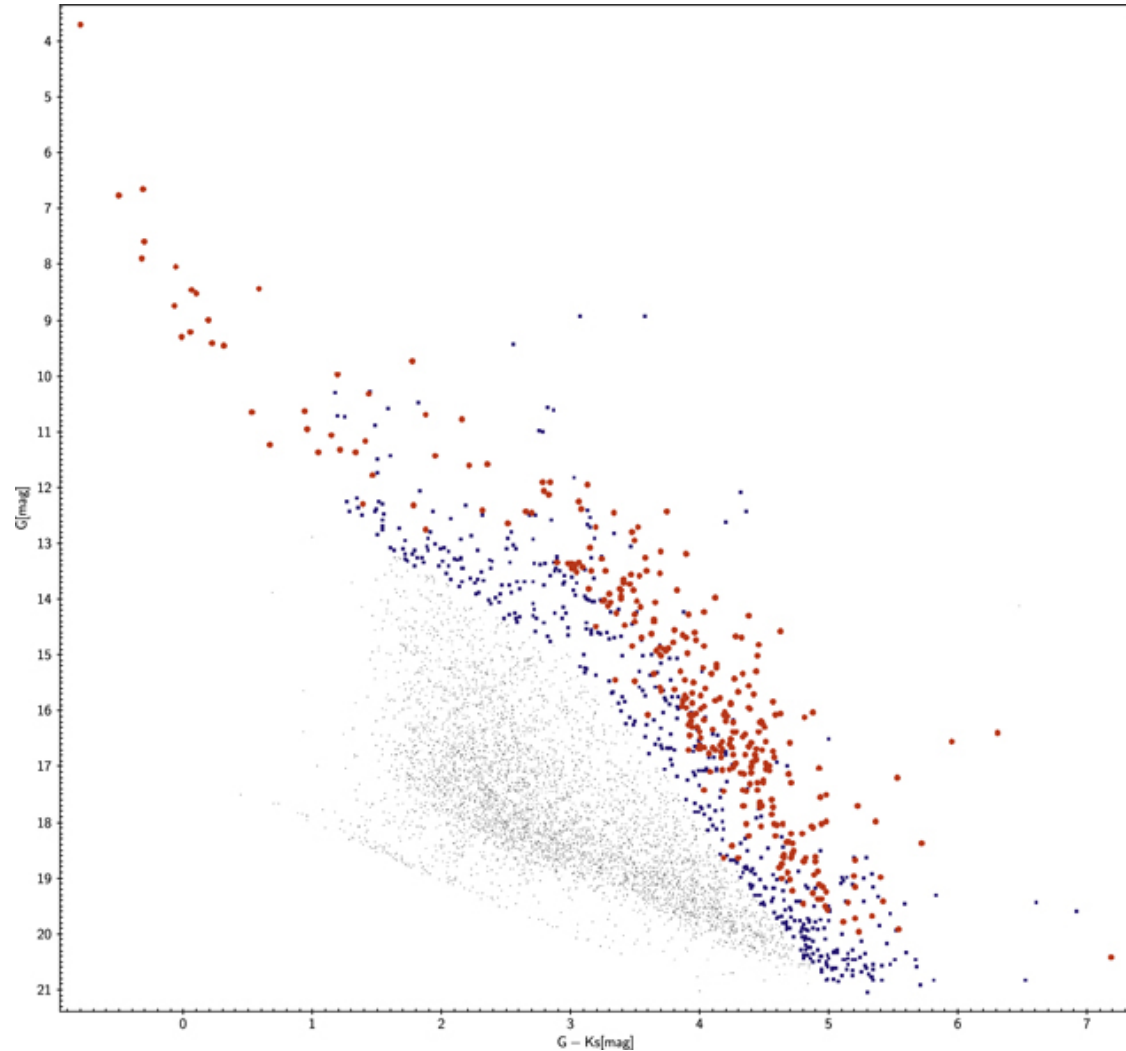


# Gaia DR2 & $\sigma$ Orionis

Colour-magnitude diagram from  $17 M_{\text{sol}}$  to  $\sim 0.05 M_{\text{sol}}$

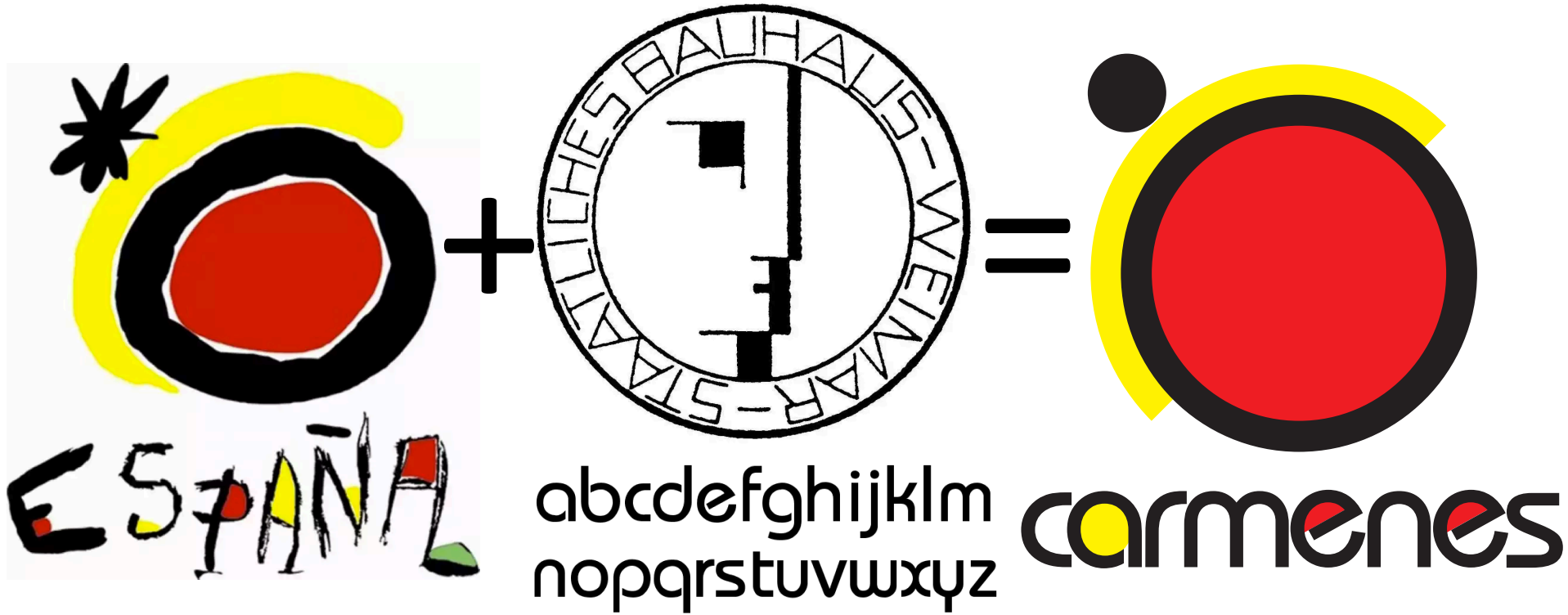
$d = 391 \pm 3$  pc, discard numerous AFG-type stars with **discordant proper motion and parallax**  $\rightarrow$  Impact on IMF (and star-to-brown dwarf ratio, disc frequency, spatial distribution...)

Caballero  
2018RNAAS...2...25C





# carmenes



MAX-PLANCK-GESELLSCHAFT

CSIC



Unión Europea  
Fondo Europeo de Desarrollo Regional  
"Una manera de hacer Europa"

Calar Alto



**C**alar  
**A**lto high-  
**R**esolution search for  
**M**dwarfs with  
**E**xoearths with  
**N**ear-infrared and visible  
**E**chelle  
**S**pectrographs

**← Calar Alto...**

3.5 m Zeiss Telescope  
2168 m, +37.2236° N  
Almería, Spain

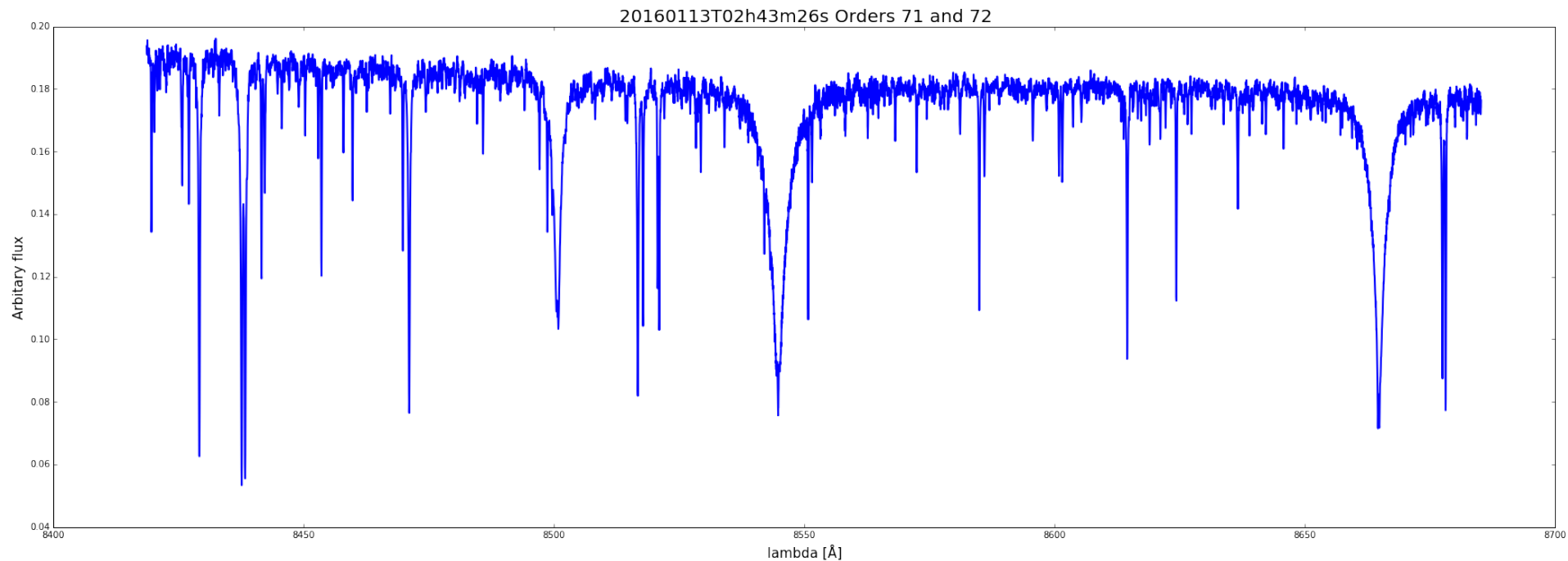


**C**alar  
**A**lto high-  
**R**esolution search for  
**M** dwarfs with  
**E**xoearths with  
**N**ear-infrared and visible  
**E**chelle  
**S**pectrographs

↓ **high-Resolution...**

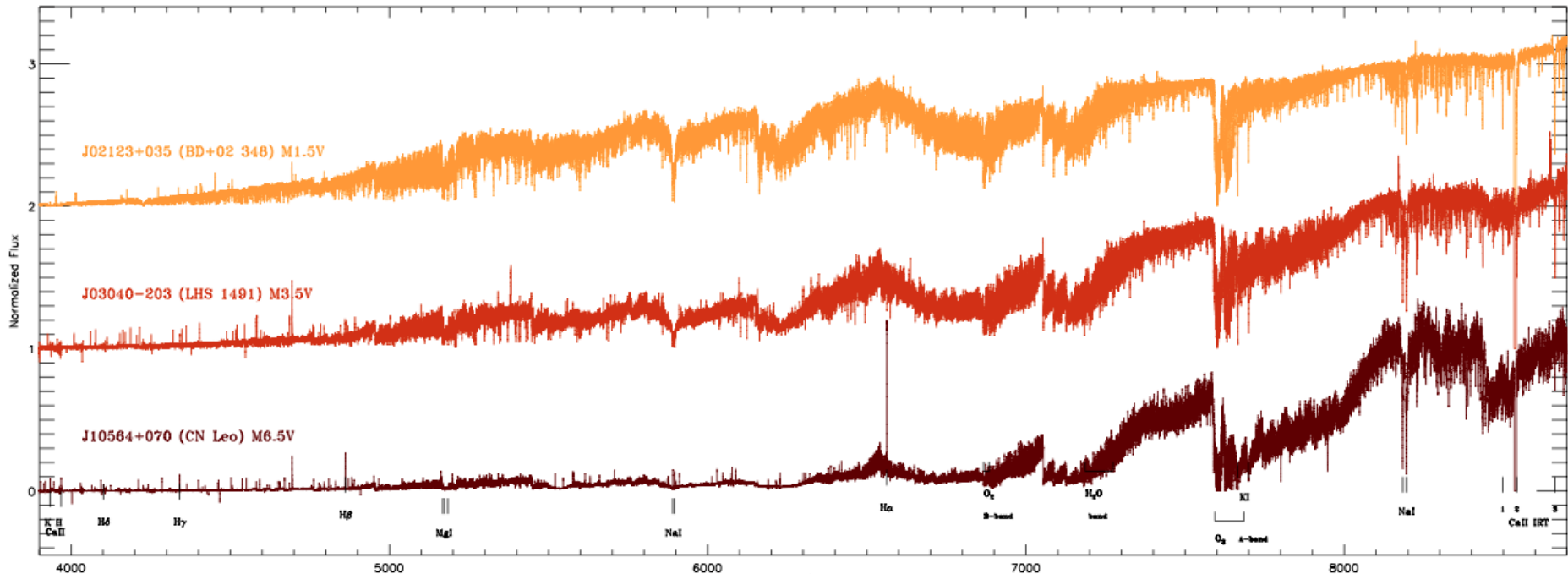
$R (= \lambda/\Delta\lambda) = 80,400 - 94,600$

Example: Ca IRT  $\lambda 850-866$  nm  
(*Gaia*, RAVE)



**C**alar  
**A**lto high-  
**R**esolution search for  
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**E**xoearths with  
**N**ear-infrared and visible  
**E**chelle  
**S**pectrographs

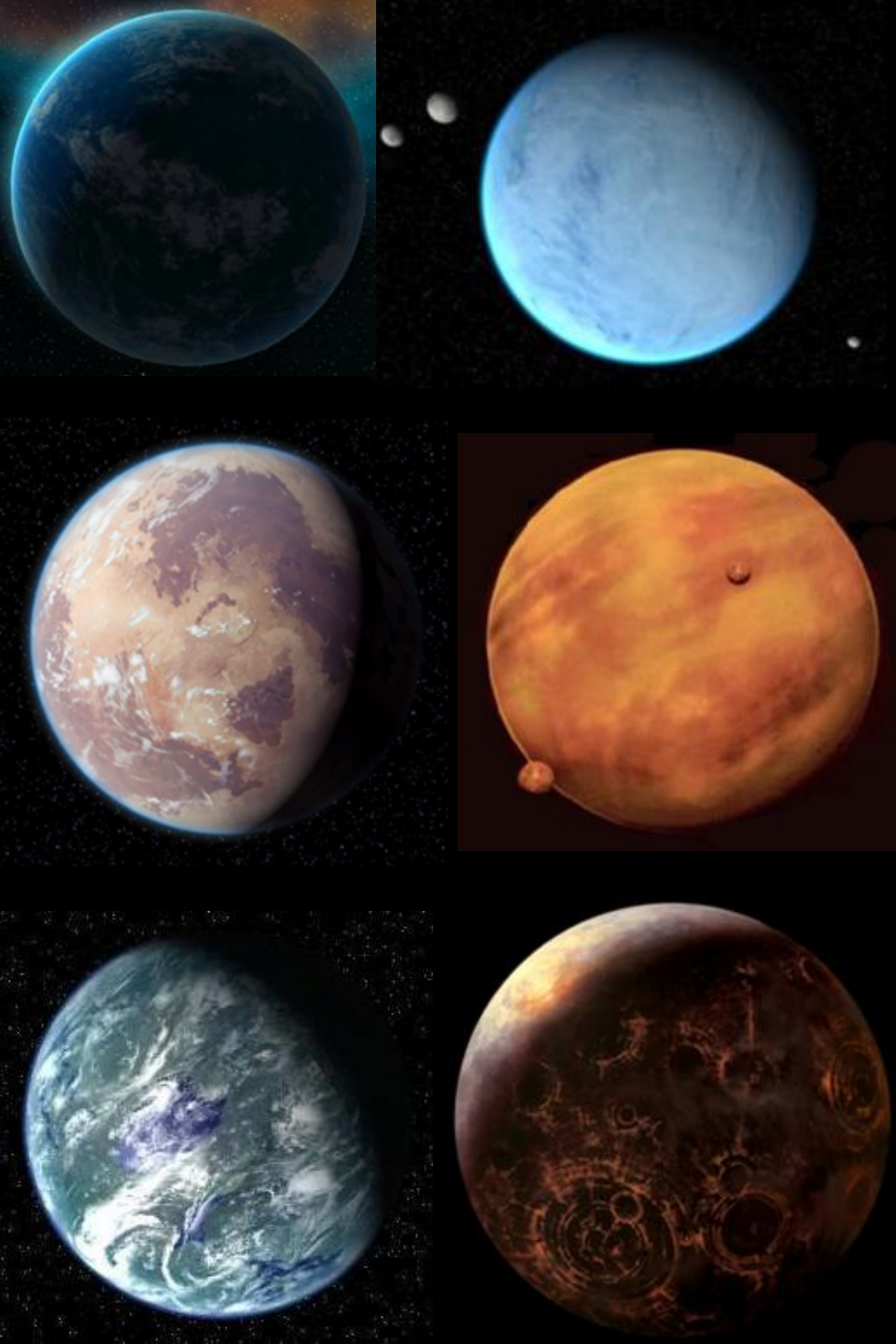
↓ **search for M dwarfs with...**  
 M0.0-9.5V, 3900-2200 K  
 Barnard's, Luyten's, vB 8



**C**alar  
**A**lto high-  
**R**esolution search for  
**M**dwarfs with  
**E**xoearths with  
**N**ear-infrared and visible  
**E**chelle  
**S**pectrographs

**← Exoearths with...**

$M = 0.8-2.0 M_{\text{Earth}}$ ,  $T_{\text{eq}} = 0-80^{\circ}\text{C}$  (classical HZ definition  
– EUV, flare)



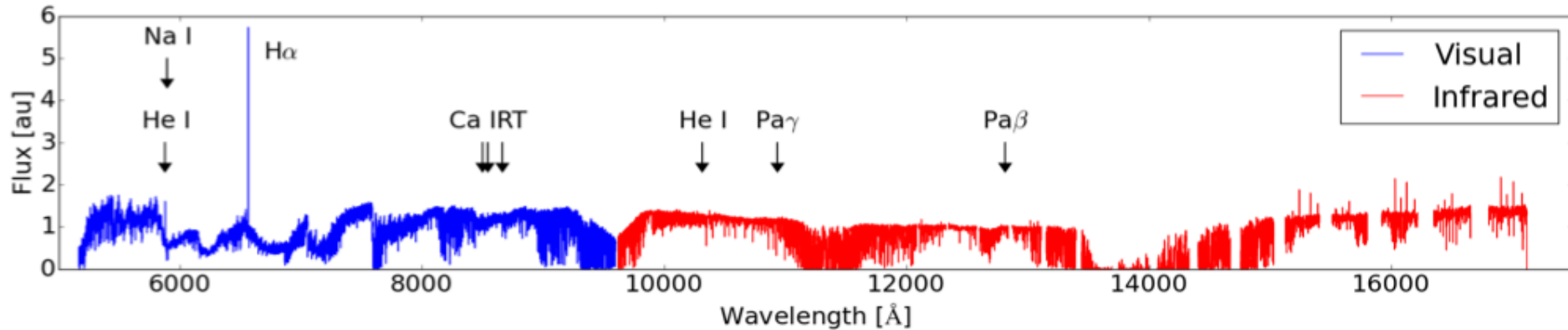


**C**alar  
**A**lto high-  
**R**esolution search for  
**M** dwarfs with  
**E**xoearths with  
**N**ear-infrared and visible  
**E**chelle  
**S**pectrographs

↓ **Near-infrared and visible...**

VIS: 0.52-0.96  $\mu\text{m}$ , e2v CCD231-84

NIR: 0.96-1.71  $\mu\text{m}$ , 2x Hawaii-2RG

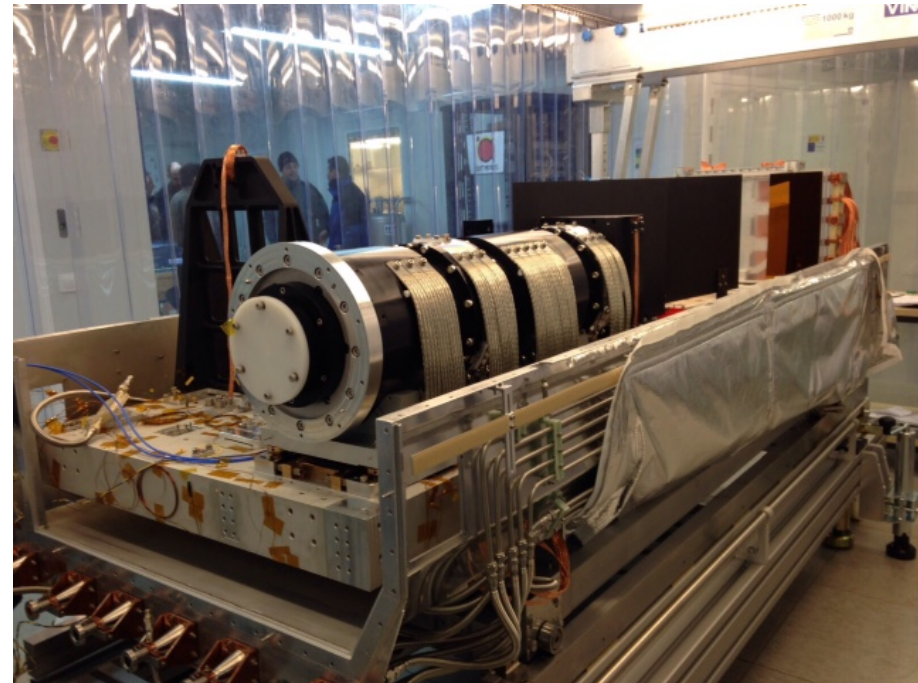
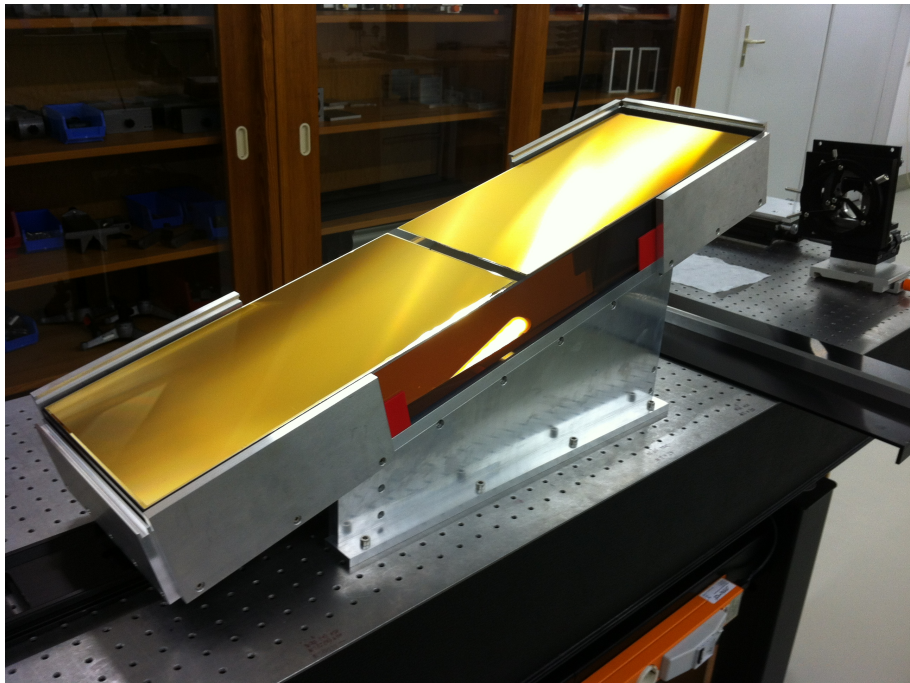


(YZ CMi, M4.5Ve)

↓ **Echelle...**

2 x Richardson Gratings R4  
( $31.6 \text{ mm}^{-1}$ ) each channel

**C**alar  
**A**lto high-  
**R**esolution search for  
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**E**xoearths with  
**N**ear-infrared and visible  
**E**chelle  
**S**pectrographs



↓ Spectrographs.

VIS:  $285.000 \pm 0.005$  K

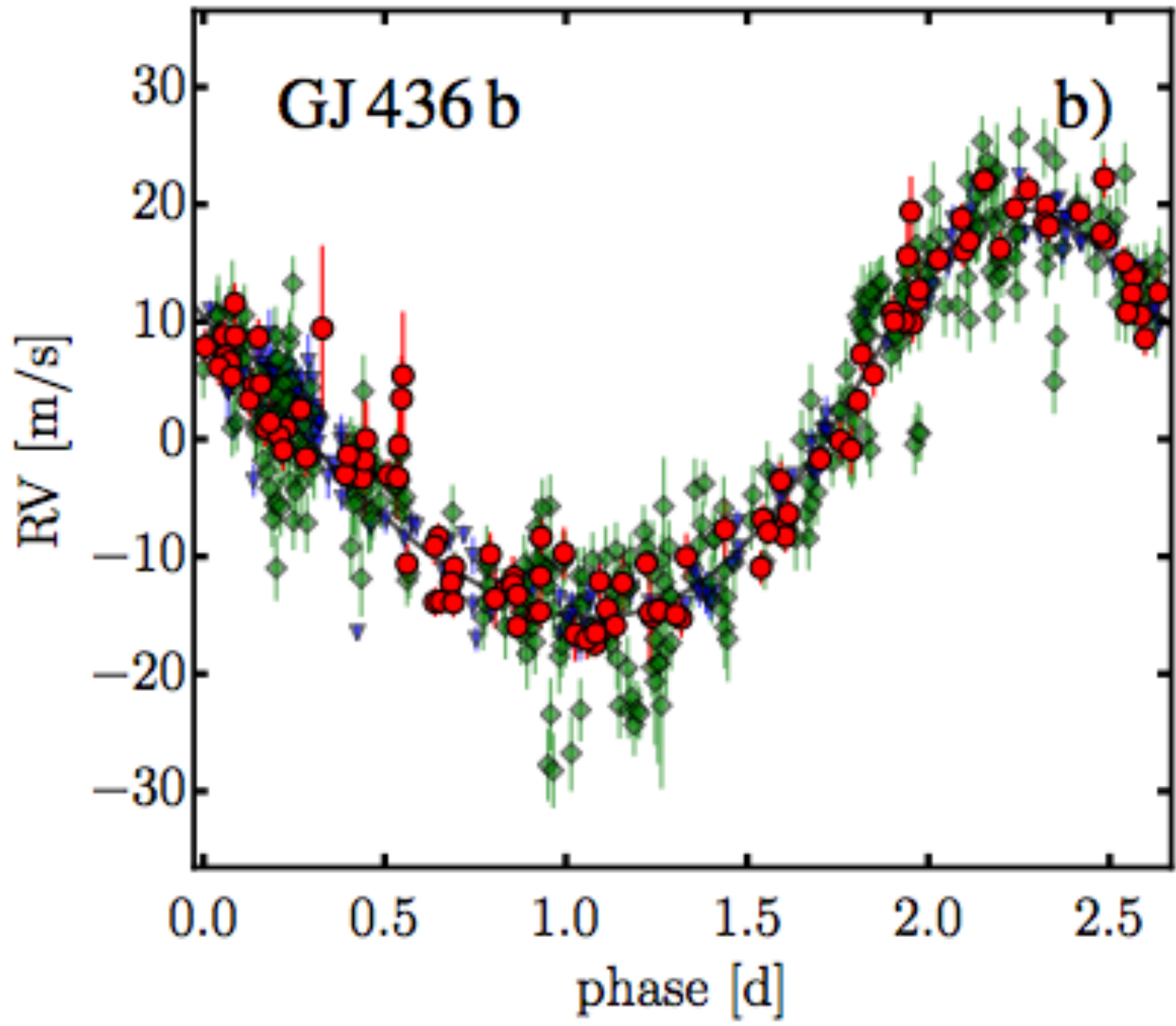
NIR:  $140.000 \pm 0.005$  K

Fibre-fed in coudé room

**C**alar  
**A**lto high-  
**R**esolution search for  
**M**dwarfs with  
**E**xoearths with  
**N**ear-infrared and visible  
**E**chelle  
**S**pectrographs









# Carmencita



Carmencita (1894)



Herbs, species and... paella!

# CARMENCITA: CARMENES Cool dwarf Information and data Archive



CARMENES input catalogue  
of **~2200 nearby bright M  
dwarfs** ( $\delta > -23$  deg)

**~300** GTO targets



Table Browser for 1: carmencita.084.csv

|    | #Karmn      | Comp | Flags | SS | Name                  | GJ     | SpT        | Ref01 | RA_J2016    | DE_J2016    |
|----|-------------|------|-------|----|-----------------------|--------|------------|-------|-------------|-------------|
| 1  | J00012+139N | Aab  | B.F   | SX | BD+13 5195            |        | M0.5 V+    | JA74  | 00:01:13.22 | +13:58:32.6 |
| 2  | J00012+139S | B    | ..F   |    | BD+13 5195B           |        | M0.0 V     | Lep13 | 00:01:12.89 | +13:58:22.0 |
| 3  | J00026+383  | -    | ..F   |    | 2MJ00024011+3821453   |        | M4.0 V     | Fri13 | 00:02:40.02 | +38:21:44.8 |
| 4  | J00033+046  | -    | ..F   |    | StKM 1-2199           |        | M1.5 V     | Lep13 | 00:03:18.97 | +04:41:11.5 |
| 5  | J00051+457  | C    |       | S3 | GJ 2                  | 2      | M1.0 V     | PMSU  | 00:05:12.18 | +45:47:09.1 |
| 6  | J00056+458  | B    | J..   | SX | HD 388                | 4B     | M0.0 V     | Cve12 | 00:05:42.31 | +45:48:34.9 |
| 7  | J00067-075  | -    |       | S1 | GJ 1002               | 1002   | M5.5 V     | PMSU  | 00:06:42.34 | -07:32:46.3 |
| 8  | J00077+603  | AB   | B..   | SX | G 217-032             |        | M4.0 V+    | AF15a | 00:07:43.42 | +60:22:53.8 |
| 9  | J00078+676  | -    | ..F   |    | 2MJ00075079+6736255   |        | M2.0 V     | Lep13 | 00:07:50.68 | +67:36:23.9 |
| 10 | J00079+080  | -    | ..F   |    | LP 524-065            | 3007   | M3.0 V     | PMSU  | 00:07:58.67 | +08:00:12.1 |
| 11 | J00081+479  | AB   | S..   | SX | 1RXS J000806.3+475659 |        | M4.0 V+    | Lep13 | 00:08:06.23 | +47:57:02.6 |
| 12 | J00084+174  | -    | ..F   |    | MCC 351               | 3008   | M0.0 V     | PMSU  | 00:08:27.19 | +17:25:26.4 |
| 13 | J00088+208  | AB   | B..   | SX | LP 404-033            | 3010   | M4.5 V+    | PMSU  | 00:08:53.60 | +20:50:20.7 |
| 14 | J00110+052  | -    | ..F   |    | G 031-029             |        | M1.0 V     | Lep13 | 00:11:04.89 | +05:12:33.1 |
| 15 | J00115+591  | -    | ..F   |    | LSR J0011+5908        |        | M5.5 V     | AF15a | 00:11:29.84 | +59:08:20.4 |
| 16 | J00118+229  | -    | ..F   |    | LP 348-040            |        | M3.5 V     | AF15a | 00:11:53.19 | +22:59:00.9 |
| 17 | J00119+330  | -    | ..F   |    | G 130-053             |        | M3.5 V     | AF15a | 00:11:55.80 | +33:03:11.1 |
| 18 | J00122+304  | -    | ..F   |    | 2MJ00121341+3028443   |        | M4.5 V     | AF15a | 00:12:13.48 | +30:28:43.9 |
| 19 | J00131+703  | -    | ..f   | S3 | TYC 4298-613-1        |        | M2.0 V     | Lep13 | 00:13:11.70 | +70:23:55.1 |
| 20 | J00132+693  | AB   | B..   | SX | GJ 11 AB              | 11AB   | M3.0 V+    | PMSU  | 00:13:18.09 | +69:19:32.3 |
| 21 | J00133+275  | -    | ..F   |    | 2MJ00131951+2733310   |        | M4.5 V     | AF15a | 00:13:19.54 | +27:33:29.1 |
| 22 | J00136+806  | A    | ..f   | S3 | G 242-048             | 3014   | M1.5 V     | AF15a | 00:13:40.56 | +80:40:00.0 |
| 23 | J00137+806  | B    | V.F   | SX | LP 012-304            | 3015   | M5.0 V     | PMSU  | 00:13:44.83 | +80:39:52.4 |
| 24 | J00154-161  | AB   | Bm.   | SX | GJ 1005 AB            | 1005AB | M4.0 V+    | PMSU  | 00:15:28.73 | -16:08:11.3 |
| 25 | J00156+722  | -    | ..F   |    | G 242-049             |        | M2.0 V     | Lep13 | 00:15:37.67 | +72:17:03.8 |
| 26 | J00158+135  | -    | ..f   | S2 | GJ 12                 | 12     | M3.0 V     | PMSU  | 00:15:49.91 | +13:33:27.4 |
| 27 | J00159-166  | AB   | B..   | SX | BPS CS 31060-0015     |        | M4.1 V+    | Shk09 | 00:15:57.93 | -16:36:57.8 |
| 28 | J00162+198E | B    |       | S1 | LP 404-062            | 1006B  | M4.0 V     | AF15a | 00:16:16.94 | +19:51:38.6 |
| 29 | J00162+198W | Aab  | S..   | SX | EZ Psc                | 1006A  | M4.0 V6+   | AF15a | 00:16:15.49 | +19:51:24.7 |
| 30 | J00169+051  | -    | ..F   |    | GJ 1007               | 1007   | M4.5 V     | PMSU  | 00:16:56.19 | +05:07:15.7 |
| 31 | J00169+200  | AB   | B.F   | SX | G 131-047             | 3022   | M3.5 V+    | PMSU  | 00:16:57.06 | +20:03:55.4 |
| 32 | J00173+291  | -    | ..f   | S3 | Ross 680              | 3023   | M2.0 V     | PMSU  | 00:17:21.09 | +29:11:05.0 |
| 33 | J00176-086  | -    | ..F   |    | BD-09 40              | 3025   | M0.0 V     | PMSU  | 00:17:41.23 | -08:40:55.6 |
| 34 | J00179+209  | A    | ..F   |    | LP 404-081            |        | M1.0 V     | Lep13 | 00:17:58.89 | +20:57:18.8 |
| 35 | J00182+102  | -    | ..f   | S3 | GJ 16                 | 16     | M1.5 V     | PMSU  | 00:18:16.59 | +10:12:09.6 |
| 36 | J00183+440  | A    |       | S3 | GX And                | 15A    | M1.0 V     | AF15a | 00:18:27.07 | +44:01:29.1 |
| 37 | J00184+440  | B    |       | S2 | GQ And                | 15B    | M3.5 V     | PMSU  | 00:18:30.00 | +44:01:44.5 |
| 38 | J00188+278  | -    | ..F   |    | LP 292-066            | 3027   | M4.0 V     | PMSU  | 00:18:54.02 | +27:48:48.1 |
| 39 | J00201-170  | -    | .mF   |    | LP 764-108            | 2003   | M1.0 V(k)+ | Gra06 | 00:20:08.54 | -17:03:41.1 |
| 40 | J00204+330  | -    | ..F   |    | LP 292-067            | 3028   | M5.5 V     | PMSU  | 00:20:30.66 | +33:04:53.6 |

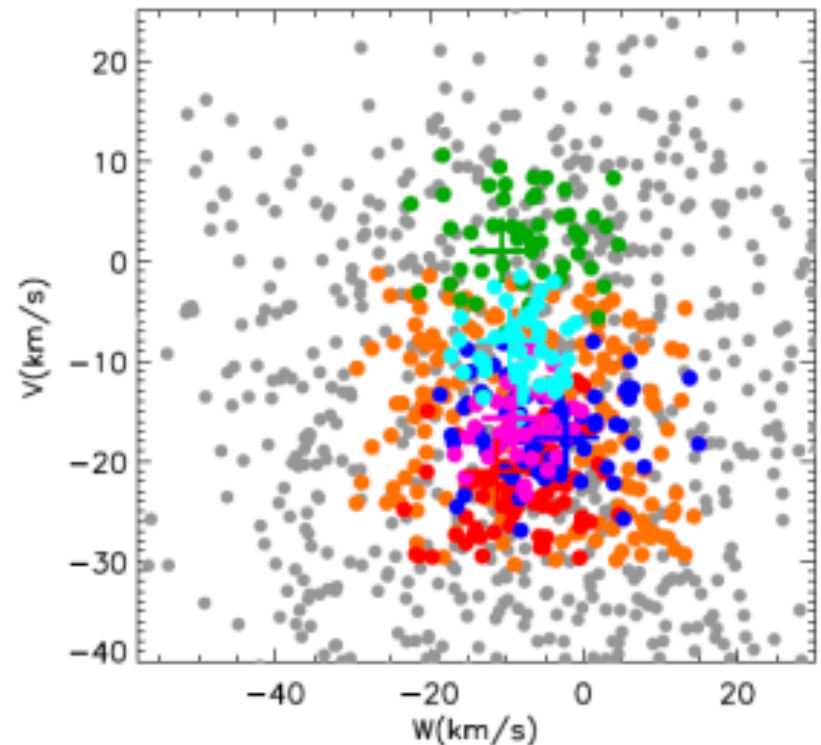
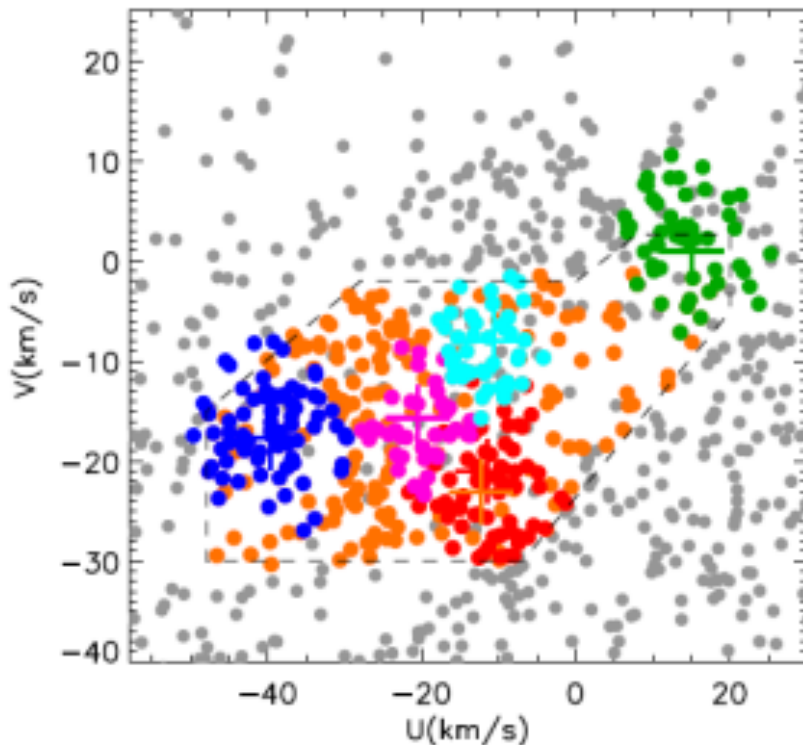
# Carmencita legacy (2200 stars)



Karmn | Comp | Flags | SS | Name | GJ | SpT |  
**RA\_J2000** | **DE\_J2000** | **muRA\_masa-1** | **muDE\_masa-1** |  
**Vr\_kms-1** | **pi\_mas** | **d\_pc** | **U\_kms-1** | **V\_kms-1** |  
**W\_kms-1** | FUV\_mag | NUV\_mag | u\_mag | BT\_mag |  
**BP\_mag** | B\_mag | g\_mag | **G\_mag** | VT\_mag |  
V\_mag | **RP\_mag** | Ra\_mag | r\_mag | i\_mag | z\_mag  
| IN\_mag | J\_mag | H\_mag | Ks\_mag | W1\_mag |  
W2\_mag | W3\_mag | W4\_mag | WideCompanion |  
WideWDS | Widerho\_arcsec | WideCompanionSpT |  
WideCompanionJ\_mag | WideCompanionFeH |  
CloseMultiplicity | CloseWDS | Closerho\_arcsec |  
pEWHalpha\_A | 1RXS | CRT\_s-1 | HR1 | HR2 | vsini\_kms-  
1 | P\_d | Flare | **Populartion** | **MovingGroup** | Teff\_K  
| logg | **R\_Rsol** | **L\_Lsol** | **M\_Msol** | Metallicity |  
Age\_Ga | LoRes\_spectrum | HiRes\_spectrum |  
LoRes\_image | HiRes\_image | RV | Planet | Origin |  
Class | Notes

# More Carmencita:

- Wide multiplicity, FGK primaries + M secondaries → Metallicity (Montes et al.)
- Wide multiplicity, white dwarfs + M...
- UVW space velocities → Young moving groups → Age! (planet formation)

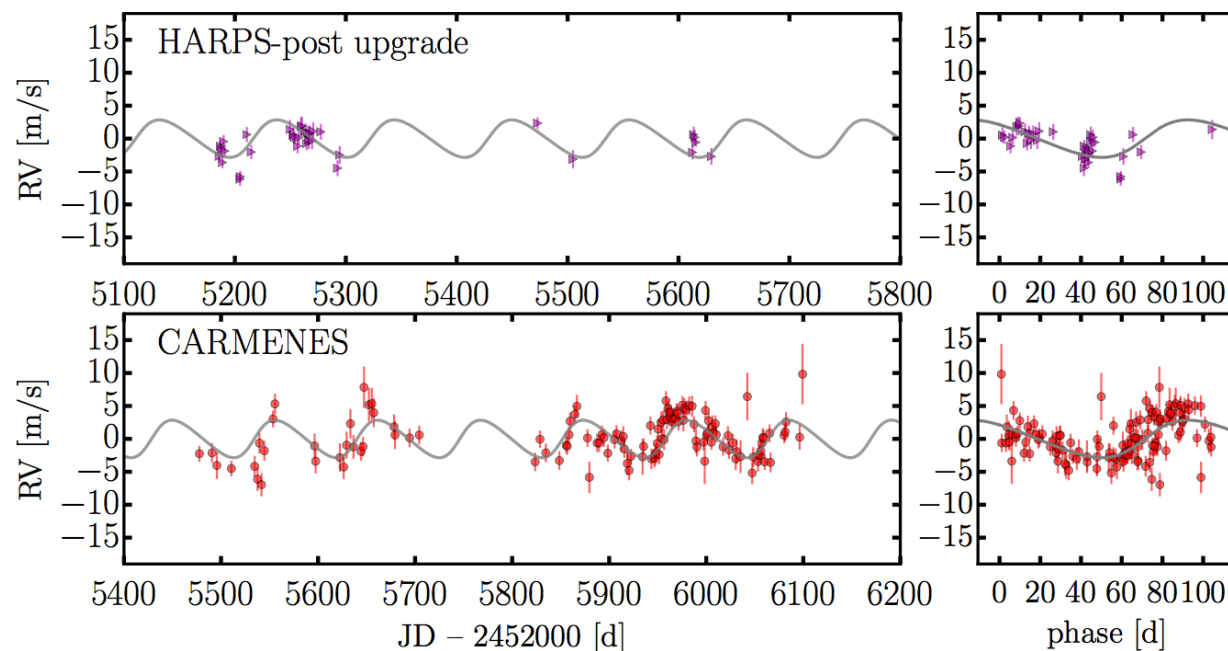
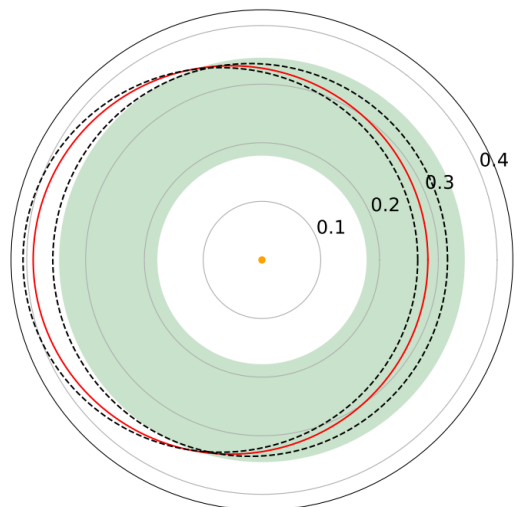


## The CARMENES search for exoplanets around M dwarfs

### A Neptune-mass planet in the habitable zone around HD




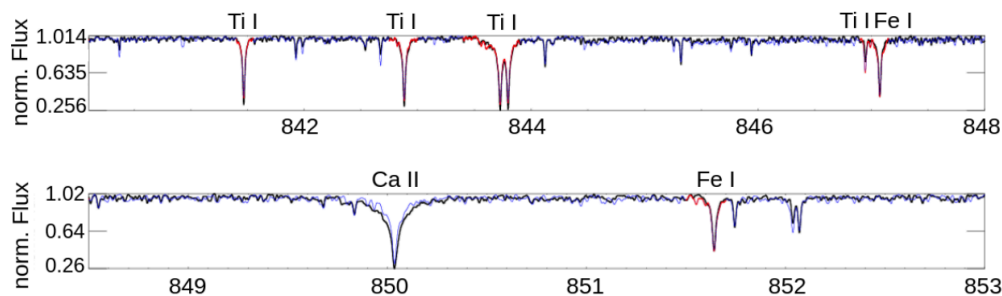
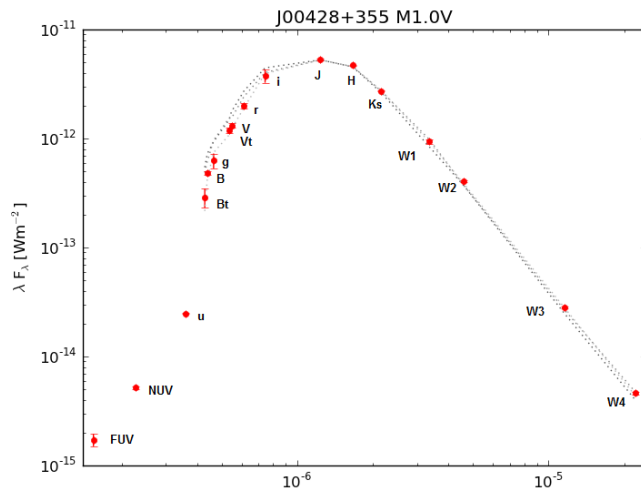
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# Stellar parameters of exoplanet hosts

Table 1: Basic information on the host star.

|   |  | Ref. <sup>a</sup> |
|---|---|-------------------|
| Karmn <sup>b</sup>                              |   |                   |
| Wolf  |   |                   |
| GJ  |   |                   |
| BD  |   |                   |
| Var. name                                       |   |                   |
| Sp. type  | M2.5 V  | AF15              |
| $G$ [mag]                                       | $8.0976 \pm 0.011$  | <i>Gaia</i>       |
| $J$ [mag]                                       | $5.583 \pm 0.030$   | 2MASS             |
| $d$ [pc]  | $5.9116 \pm 0.018$  | <i>Gaia</i>       |
| $\mu_\alpha \cos \delta$ [mas a <sup>-1</sup> ] | $-579.043 \pm 0.088^c$  | <i>Gaia</i>       |
| $\mu_\delta$ [mas a <sup>-1</sup> ]             | $-1332.743 \pm 0.081^c$   | <i>Gaia</i>       |
| $V_r$ [km s <sup>-1</sup> ]                     | +35.678   | Rei18             |
| $U$ [km s <sup>-1</sup> ]                       | +53.2   | This work         |
| $V$ [km s <sup>-1</sup> ]                       | -7.6  | This work         |
| $W$ [km s <sup>-1</sup> ]                       | -5.0  | This work         |
| $v \sin i$ [km s <sup>-1</sup> ]                | <2  | Rei18             |
| $T_{\text{eff}}$ [K]                            | $3557 \pm 51$   | Pas18             |
| $\log g$  | $4.86 \pm 0.07$   | Pas18             |
| [Fe/H]  | $0.00 \pm 0.16$   | Pas18             |
| $L$ [L <sub>⊙</sub> ]                           | $0.0296 \pm 0.0007$   | This work         |
| $R$ [R <sub>⊙</sub> ]                           | $0.453 \pm 0.019$   | This work         |
| $M_\star$ [M <sub>⊙</sub> ]                     | $0.45 \pm 0.04$   | This work         |
| pEW(H $\alpha$ ) [Å]                            | +0.3±0.1  | Jef18             |
| $P_{\text{rot}}$ [d]                            | $46.04 \pm 0.20$  | DA17              |
| Kinematic pop.                                  | Thin disc   | CC16              |

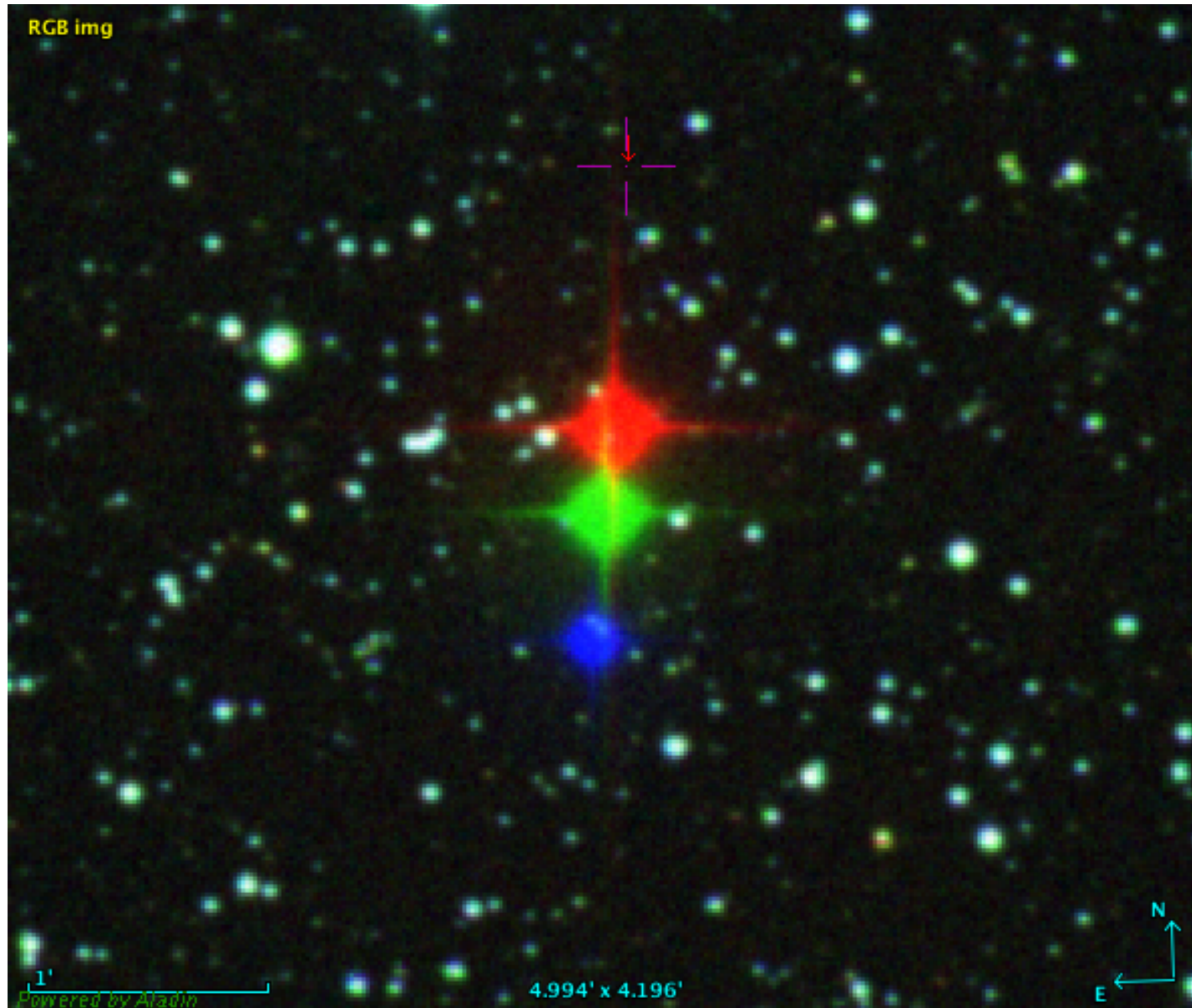


$$L = 4 \pi R^2 \sigma T_{\text{eff}}^4$$

$$M_\star - R$$



And many more...



A detailed view of the Gaia satellite in space, showing its large white sunshield and the main body of the spacecraft. The background is a deep blue and black space filled with numerous stars of various colors and sizes. The satellite's structure is complex, with various panels and components visible. The sunshield is a large, white, multi-layered structure that covers the main body of the satellite. The main body is a cylindrical structure with various instruments and sensors. The overall scene is a high-quality digital rendering of the Gaia mission in space.

# Gaia DR1

*A soundtrack for the ESA billion star surveyor*  
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