



¿Que puede aportar VO a proyectos grandes como el Gaia-ESO Survey y a futuros proyectos de la REG?

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REG Working Groups

- G1: [Estrellas masivas y distancias exactas a cúmulos masivos](#)
- G2: [Cinturon de Gould](#)
- G3: [Cúmulos estelares y asociaciones](#)
- G4: [Evolución estelar \(tardía\)](#)
- G5: [Estructura de las estrellas y comparación con modelos](#)
- G6: [Análisis estadístico de las poblaciones estelares](#)
- G7: [Estrellas de baja masa, enanas marrones y exoplanetas](#)
- G8: [Estructura, cinemática y dinámica a gran escala](#)
- G9: [Instrumentación](#)
- G10: [Sistema Solar](#)
- G11: [Observatorio Virtual](#)
- G12: [Minería de datos y astroestadística](#)

- **VO's main goal:** *“Efficient access and analysis of the information hosted in astronomical archives.”*

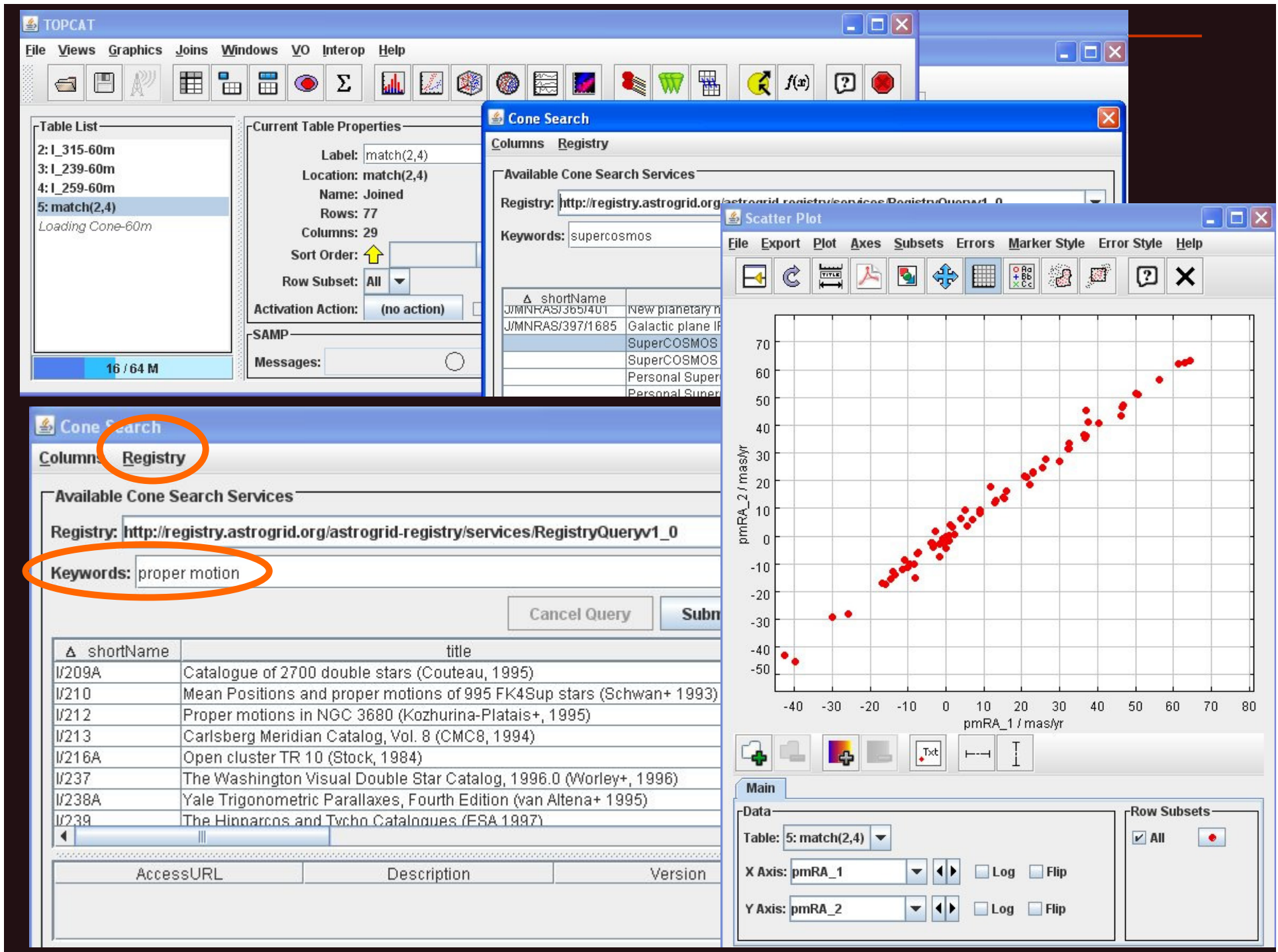
GAIA-ESO survey

- Public spectroscopic survey ($> 100\,000$ stars)
- Field stars and clusters
- From halo to star forming regions.
- First homogeneous overview of kinematics and abundances.
- FLAMES / GIRAFFE-UVES
- 250 scientists
- 300 nights. P88 - P97

Gaia-ESO Survey WG4: Cluster Stars Target Selection

- The clusters to be observed
- The stars to be observed in each cluster.
- Identification of the best candidates by looking through archives.
 - What is already available?
 - Cluster: distance, reddening, metallicity,...
 - Stars: photometry, spectroscopy,...

Virtual Observatory as a discovery tool



A typical VO workflow

Virtual Observatory as an analysis tool

■ NGC 1817 (Balaguer & Jordi)

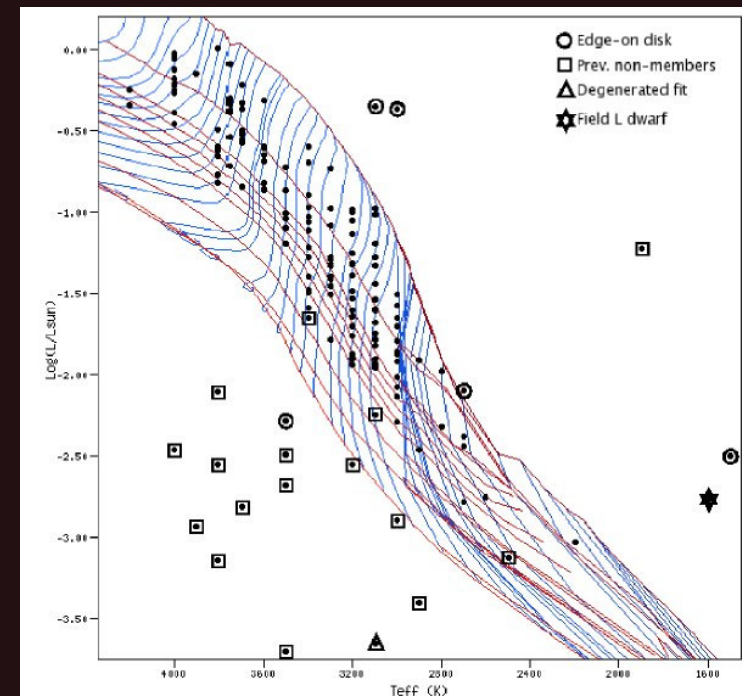
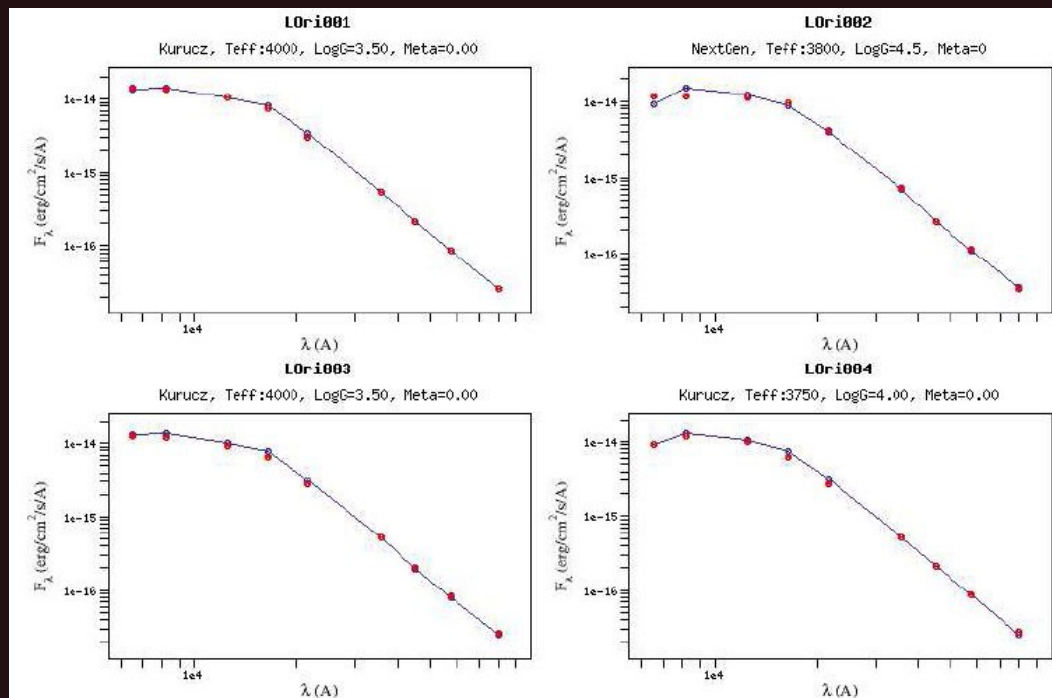
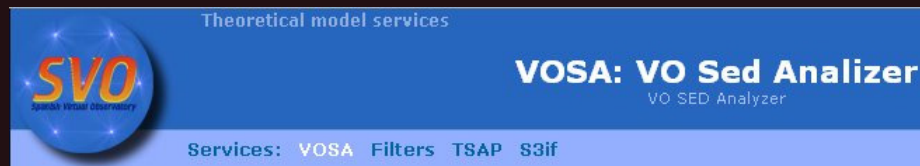
- *“Given a cone of the sky, give me all stars with photometry in SDSS, 2MASS and Balaguer & Núñez (2004)”.*



- `java -jar ~esm/stilts.jar tpipe in='http://vizier.u-strasbg.fr/viz bin/votable/-A?-source=II/294&RA=78.025&DEC=16.7&SR=0.5' out=sdss.xml ofmt=votable`
- `java -jar ~esm/stilts.jar tskymatch2 ifmt1=votable in1=sdss.xml in2=2mass.xml ra1="RAJ2000" dec1="DEJ2000" ra2="RAJ2000" dec2="DEJ2000" error=2 find=best out=cross.xml ofmt=votable`

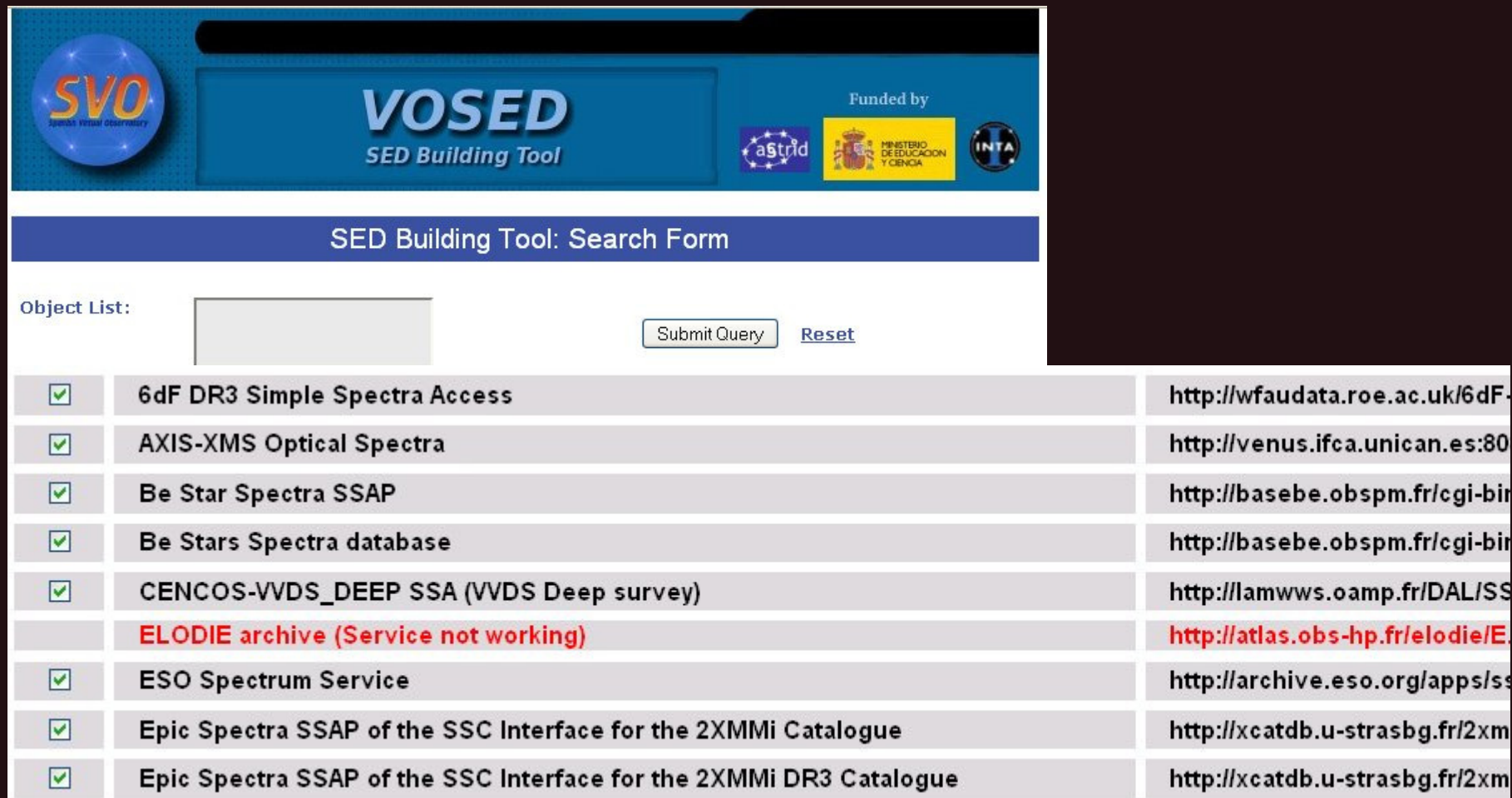
A typical VO workflow: NGC1817

- “And I want to complement this photometry with other available in archives and derive physical parameters (e.g. T_{eff}) and build and H-R diagram.”



A typical VO workflow: NGC1817

- “And I want to get all the spectroscopic information that is available in VO archives.”



SED Building Tool: Search Form

Object List: [Reset](#)

☒	6dF DR3 Simple Spectra Access	http://wfaudata.roe.ac.uk/6dF-
☒	AXIS-XMS Optical Spectra	http://venus.ifca.unican.es:80
☒	Be Star Spectra SSAP	http://basebe.obspm.fr/cgi-bin
☒	Be Stars Spectra database	http://basebe.obspm.fr/cgi-bin
☒	CENCOS-VVDS_DEEP SSA (VVDS Deep survey)	http://lamwww.oamp.fr/DAL/SS
☐	ELODIE archive (Service not working)	http://atlas.obs-hp.fr/elodie/E
☒	ESO Spectrum Service	http://archive.eso.org/apps/ss
☒	Epic Spectra SSAP of the SSC Interface for the 2XMMi Catalogue	http://xcatdb.u-strasbg.fr/2xm
☒	Epic Spectra SSAP of the SSC Interface for the 2XMMi DR3 Catalogue	http://xcatdb.u-strasbg.fr/2xm

Gaia-ESO Survey WG9: Determination of stellar parameters

■ Comparison with theoretical models

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ATLAS VERSUS NEXTGEN MODEL ATMOSPHERES: A COMBINED ANALYSIS OF SYNTHETIC SPECTRAL ENERGY DISTRIBUTIONS

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² See <http://kurucz.harvard.edu>.

³ See <ftp://calvin.physast.uga.edu/pub/> and <http://dilbert.physast.uga.edu/~yeti>. Note that the libraries of dwarf and giant stars available at these sites have lower T_{eff} limits than the published ones.

Accessing Kurucz & Nextgen models

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¹ Available via anonymous FTP from <ftp://calvin.physast.uga.edu/pub/NextGen> or via the WWW URL <http://dilbert.physast.uga.edu/~yeti>.

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[GRIDM05](#)

[GRIDM050DFNEW](#)

[ddop00k8.dat19](#)

13-Aug-2009

9 11:40 623K
9 11:40 55K
9 11:40 709K
13-Aug-2009 11:40 54K
1:40 701K
1:40 240
1:40 1.6K
1:40 799
1:40 12M
1:40 11M
1:40 12M
1:40 11M
1:40 9.8M
1:40 9.7M
1:40 11M
1:40 11M
1:40 11M
1:40 11M



Servidor no encontrado

Firefox no puede encontrar el servidor en dilbert.physast.uga.edu.

- Compruebe que no ha cometido errores al escribir la dirección, como **ww**.example.com en lugar de **www**.example.com
- Si no puede cargar ninguna página, compruebe la conexión de red de su ordenador.
- Si su ordenador o su red están protegidos por un cortafuegos o un proxy, cerciórese de que se le permite acceder a la Web con Firefox.

Reintentar

should n
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[GRIDM30](#)

[GRIDM35](#)

[GRIDM40](#)

[fsun.pck13](#)

[fsun.pck19](#)

13-Aug-2009

13-Aug-2009

11:40 37K
11:40 37K

Kurucz: data characterisation

```
SDSC GRID [+0.0] VTURB 2.0 KM/S L/H 1.25
PROGRAM READFLUX
C SAMPLE PROGRAM READS THIS FILE ON UNIT 1
  DIMENSION Hnu(1221),HnuCONT(1221),WAVE(1221)
  CHARACTER*80 TITLE
  DO 11 ISKIP=1,22
11 READ(1,1)
C wavelength in nm
  READ(1,1) WAVE
  1 FORMAT(10F10.2)
  DO 8 MODEL=1,500
C ergs/cm**2/s/hz/ster
  READ(1,2,END=9) TITLE
  2 FORMAT(A80)
  PRINT 3,MODEL,TITLE
  3 FORMAT(15,1X,A80)
  READ(1,4) Hnu
  READ(1,4) HnuCONT
  4 FORMAT(8E10.4)
  8 CONTINUE
  9 CALL EXIT
  END
  9.09      9.35      9.61      9.77      9.96      10.20      10.38      10.56
10.77      11.04      11.40      11.78      12.13      12.48      12.71      12.84
13.05      13.24      13.39      13.66      13.98      14.33      14.72      15.10
15.52      15.88      16.20      16.60      17.03      17.34      17.68      18.02
18.17      18.61      19.10      19.39      19.84      20.18      20.50      21.05
21.62      21.98      22.30      22.68      23.00      23.40      24.00      24.65
```

Gaia-ESO Survey WG9: Determination of stellar parameters

Comparison with theoretical models

Granada Stellar Seismic Models

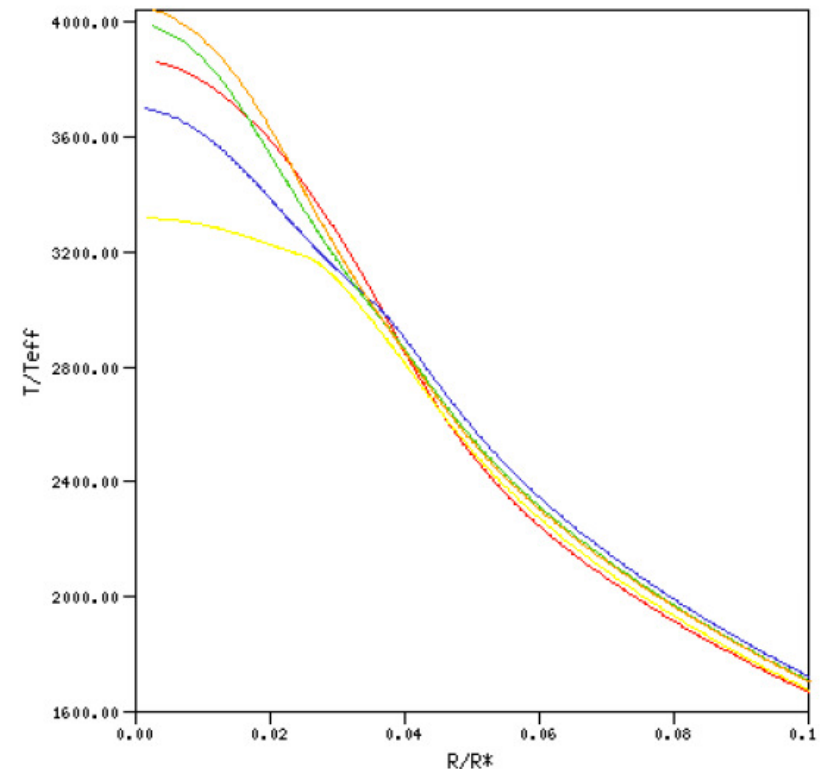
Granada Stellar Seismic Models (GSSM-VO) adapts the Granada Team numerical package outputs to be used in VO in order to perform on-line stellar seismology. This package contains the evolutionary codes [CESAM](#) and [CESAM2K](#) and two oscillation codes: [GraCo](#) and [FILOU](#).

CESAM2k evolutionary code + GraCo oscillation code

You can search the database in terms of several parameters (move your mouse over the (?) symbol to see a description and available range of values for each one).

- Please, select a range for each parameter that you want to use in the search and then click the "Search" button to get a list of the available files.
- Take into account that some combinations of values could correspond to no result.

Structure search parameters			Sismology search parameters		
(?)	Teff	4000 - 5000 (K)	(?)	F0	- (muHz)
(?)	Lum	- (Lsun)	(?)	F1	- (muHz)
(?)	Log(g)	-	(?)	F0/F1	-
(?)	Density	- (g/cm3)	(?)	$\Delta(v)$	25 - 30 (muHz)
(?)	Age	- (Myr)	$\Delta(v)$		
(?)	[Fe/H]	-	Large separation		
(?)	Z	-	$\Delta(v) = \text{freq}(n, l, m=0) - \text{freq}(n-1, l, m=0)$		
(?)	Hcent	-	The large separation is computed as an average of the average for the oscillation modes with Frequency and a mode degree		
(?)	R*	- (Rsun)	(?)	VSta	- (muHz)
(?)	Mass	- (Msun)			
(?)	Vrot	- cm/s			
(?)	Wrot	- rad/s			
(?)	Trot	- sec			
(?)	α_{MLT}	-			
(?)	Over.	-			
Search			Reset		



Archive of preparatory / follow-up data.

- Raw, calibration, reduced
- High level data products
 - Radial velocities, metallicities, equiv. widths, classification,...
- VO-compliance
 - Easy interoperability with the rest of archives.

How to undertake all these activities?

■ VO-REG

- VO support to REG scientific projects
- Technical developments:
 - Development and maintenance of VO archives and services
- Data Mining

The IDEAS Work Programme

***EUROPEAN RESEARCH COUNCIL WORK PROGRAMME
2012***