

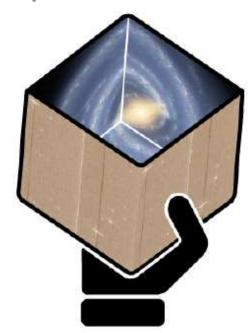






# SOBRE LA PRIMERA Y SEGUNDA PUBLICACIÓN DE DATOS

(based on a presentation by A. Brown)





**EEC** IV Reunión Científica REG – Barcelona 23-25 Mayo 2016

## http://www.cosmos.esa.int/web/gaia/release

- 1st release (summer 2016):  $(\alpha,\delta,G)$  for all well-behaved sources, variable, TGAS  $(\alpha,\delta,G,\mu,\pi)$
- 2nd release (end 2017):  $(\alpha, \delta, G, \mu, \pi, G_{BP}, G_{RP}, V_{rad})$  for all well-behaved sources
- **3rd release (2018):** Binary orbits, Classification + astrophysical parameters
- 4th release (2019): Variable classification + epoch photometry, Solar system, Non-single star catalogue.
- **Final release (2022):** Full astrometric, photometric and radial-velocity catalogues, variable stars, non-single, classifications, astrophysical parameters, exoplanet list, epoch data for all sources, ground-based observations for data-processing.

## Gaia-DR1 contents

### Astrometry

 $\alpha$ ,  $\delta$  for all sources (> 1 billion)

 $\varpi$ ,  $\mu_{\alpha*}$ ,  $\mu_{\delta}$  for TGAS sources ( $\sim 2$  million)

Covariance matrices (standard errors and correlations)

→ formal errors 'inflated' to realistic values

Statistical information to judge astrometry quality

### **Photometry**

Mean G-band fluxes and errors for all sources

G magnitudes in VEGAMAG system

Photometric zero-points for VEGAMAG and AB

No pass-band calibration, transformation to other systems will be provided

Statistical information to judge photometry quality

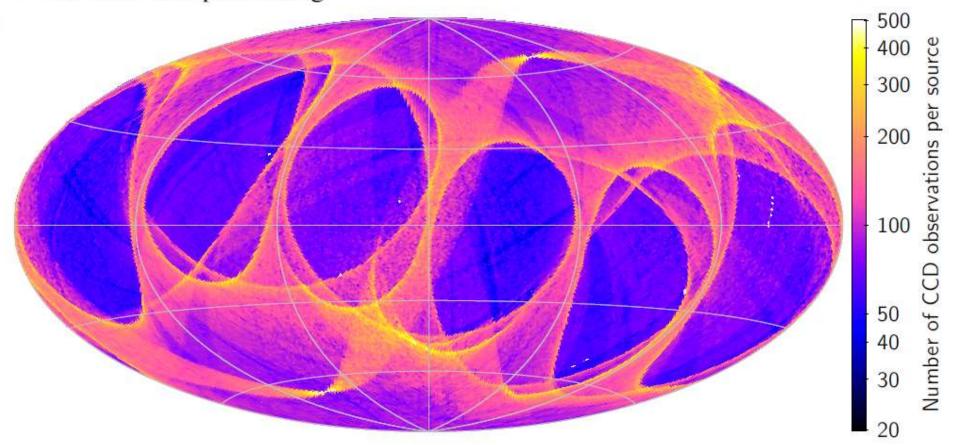
#### Variable star data

G-band light curves for selected RR-Lyrae and Cepheids

Classification information for these variables

Statistical information on G-band time series

- Gaia-DR1 results will be immensely valuable
  - most accurate sky-map to date at HST-like resolution
  - large increase in parallax information
  - much more accurate proper motions for Hipparcos stars
- Experience from scientific use of the data will benefit future improvements of the Gaia data processing



## Gaia-DR1 processing status

### General

- 14 months of input data used
- $\sim 2.3 \times 10^{10}$  transits across focal plane
- all sources treated as single
- preliminary validation of astrometry and photometry completed

### Astrometric processing

- Finished, results in main data base at ESAC
- $\bullet$  positions for > 1 billion sources ( $m \lesssim 20$ )
- parallax and proper motion for 2 million Hipparcos and Tycho-2 stars  $(m \lesssim 11.5)$
- reference frame aligned to ICRS, astrometry at epoch J2015.0
- basic angle variation correction derived from on-board metrology

## Photometric processing

- Finished, results in main data base at ESAC
- mean G-band magnitudes for all sources
- Light-curves for selected RR-Lyrae and Cepheids, mainly around ecliptic poles
  - Light-curve processing (classification and characterization) finished, data to be integrated into MDB

## Gaia-DR1: next steps

- 1. Integrate processing results into main data base (done)
- 2. Final validation of the results by independent DPAC team (CU9, ongoing)
- 3. Agreement by Gaia Science Team and DPAC Executive on Gaia-DR1 contents
  - contents of Gaia-DR1 subject to validation and GST/DPACE approval
- 4. Prepare documentation (ongoing)
  - ▶ including Gaia-DR1 papers for A&A edition
- 5. Provide details on data model and statistical information on contents
- 6. Open the Gaia archive (end of summer 2016)
- 7. Go crazy with the data...

## Gaia-DR1 contents

## Filtering before release

- Contents of Gaia-DR1 are not a 1-to-1 copy of main data base contents
- Filtering of results will be done based on validation at the AGIS, PhotPipe and VariPipe level, and at global level (CU9)
- YOUR FAVOURITE SOURCE(S) MAY THUS NOT APPEAR IN Gaia-DR1

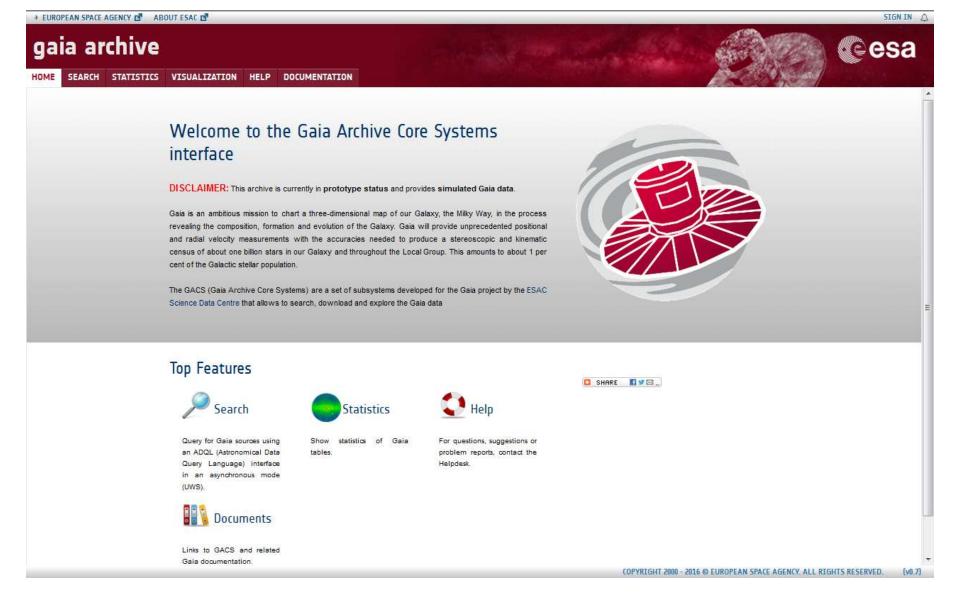
## Examples

- Omit sources with too few observations
- Omit sources without astrometry and/or photometry
- Upper limit on errors in parallax, position, photometry
- Omit sources suffering from specific problems in the data processing (e.g., data gaps)
- No high proper motion stars ( $\mu > 3.5$  arcsec yr<sup>-1</sup>) due to technical issue
- ...

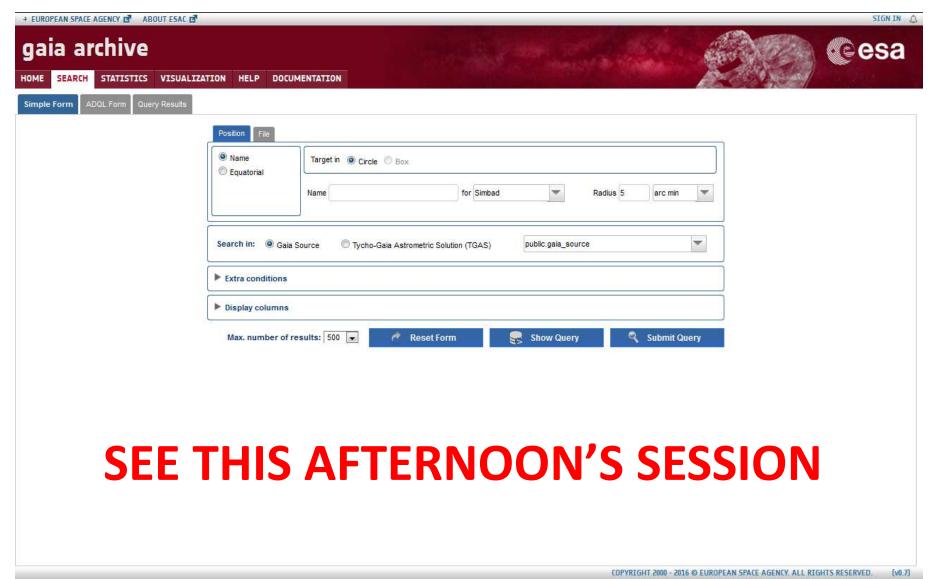
## Gaia-DR1: facilities

- Online archive @ ESDC (ESAC Science Data Centre)
  - catalogue mirrored at CDS, ASDC, ARI, AIP, as well as STScI, USNO, NAOJ, SAAO, ObsPM
  - these data centres may layer their own services on top of Gaia catalogue
- Online (interactive) and offline documentation
  - archive contents and the meaning of the tables
  - how was the processing done
  - how were the results validated
  - A&A papers providing 'condensed documentation'
  - science verification papers on open clusters and Cepheids
- Query tools for the archive (ADQL, TAP)
- Visualization tools
  - pre-computed views of large data sets
  - web-client with linked views, option to visually generate catalogue queries
- Pre-computed and validated cross-match with:
  - Hipparcos-2, Tycho-2, 2MASS PSC, GSC2.3, PPMXL, UCAC4, SDSS DR10/DR12, AllWISE, URAT-1, RAVE

# http://gaia.esac.esa.int/archive/



## http://gaia.esac.esa.int/archive/



## Plans for Gaia-DR2

- Planned for autumn 2017
  - ▶ DPAC schedule under heavy revision
- Astrometry and photometry based on roughly 22 months of data
  - Gaia stand-alone astrometric solution (unlike TGAS, no priors needed)
  - 5-parameter astrometry for all sources
- Broad band colours  $(G_{BP} G_{RP})$ 
  - improved photometric calibrations
  - proper pass-band calibrations
- Median radial velocities for bright ( $G_{RVS} < 12$ ) stars
- More variable star results
  - Cepheids, RR Lyrae all sky, LPV, short time scale variables, exercise exo-planet transit algorithms, QSO variability
- Astrophysical parameters
  - attempt determination  $T_{\text{eff}}$  and  $A_0$  from  $(G_{BP} G_{RP})$  or publish relation between  $T_{\text{eff}}$  and  $(G_{BP} G_{RP})$
  - ▶ from BP/RP spectra:  $T_{\text{eff}}$ ,  $A_0$ , [Fe/H],  $\log g$  ([Fe/H],  $\log g$  only for brighter stars)
  - $\blacktriangleright$  for TGAS stars release  $M_G$
  - attempt radius and luminosity determination for TGAS sources

## **THANK YOU**